# ATHENA

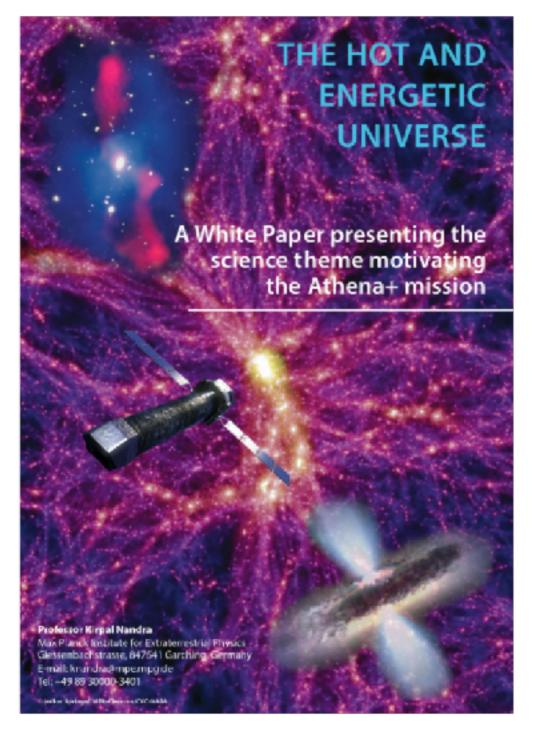
http://www.the-athena-x-ray-observatory.eu

# Cluster science with Athena

#### Etienne Pointecouteau

SnowCluster, March 23<sup>rd</sup> 2018

## The hot and energetic Universe



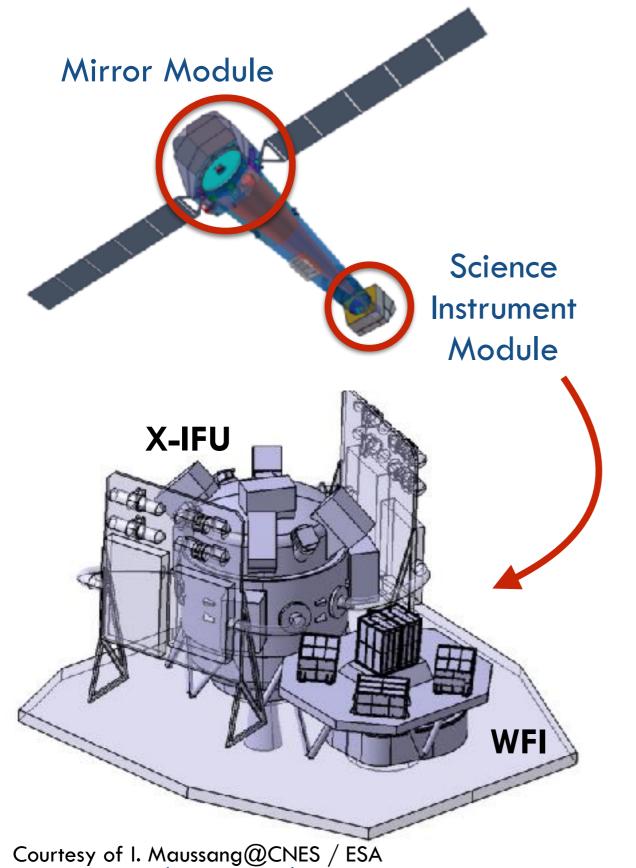
Nandra, Barret, Barcons et al. arXiv:1306.2307

- How does the ordinary matter assemble into the large-scale structures that we see today?
- How do black holes grow and influence the Universe?
- In addition:
  - Fast ToO capability to study transient sources
  - Observatory science across all corners of Astrophysics

http://www.the-athena-x-ray-observatory.eu



#### ATHENA: Advanced Telescope for High ENergy Astrophysics



- 12 m focal length
- 1.4 m<sup>2</sup> area @1 keV, 5" HEW
- Wide Field Imager (WFI)
  - Active Pixel Sensor Si detector
  - ▶ 40'x40' FoV
  - ▶ 120-150 eV @6 keV
- X-ray Integral Field Unit (X-IFU)
  - Cryogenic imaging spectrometer
  - 5' diameter equivalent FoV
  - ▶ 2.5 eV @7 keV
- Launch 2030, Ariane 6.4, L2 orbit
- 4 years nominal life-time + extensions



- Athena @ ESA
  - Cost cap = 1.05B€, encompasses the optics, payload, launch+
  - Phase A, i.e., accepted proposal in feasibility/definition phase
  - Main Milestone: Mission Adoption Review by ESA in 11/2021
- Athena optics
  - Silicon Pore Optics developed by ESA/Cosine are on a critical path
  - Maintain pressure on ESA to secure the effective area and the adequate spatial resolution
- Instruments (X-IFU and WFI)
  - Funded by national agencies + contributions by ESA, NASA, JAXA
  - Instruments (X-IFU and WFI) configuration as proposed

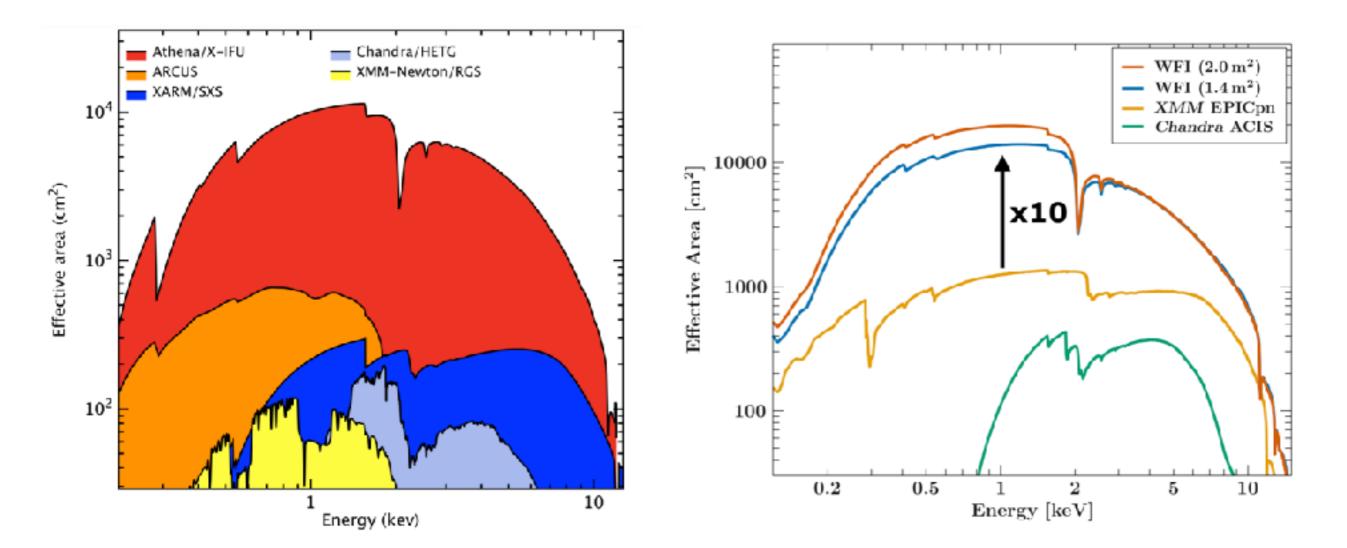
4



# Effective area

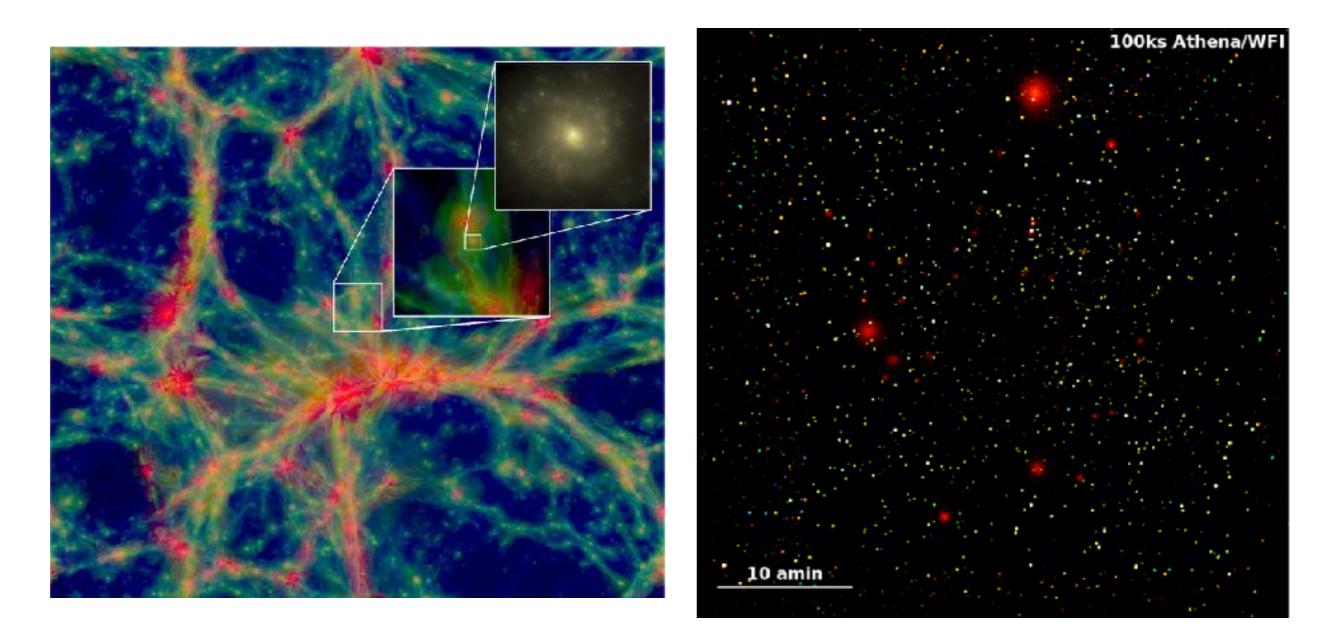
X-IFU

WFI





## Assembly of the first massive halos



Schaye et al. 2015

Athena/WFI 1Ms simulation MPE & WFI team

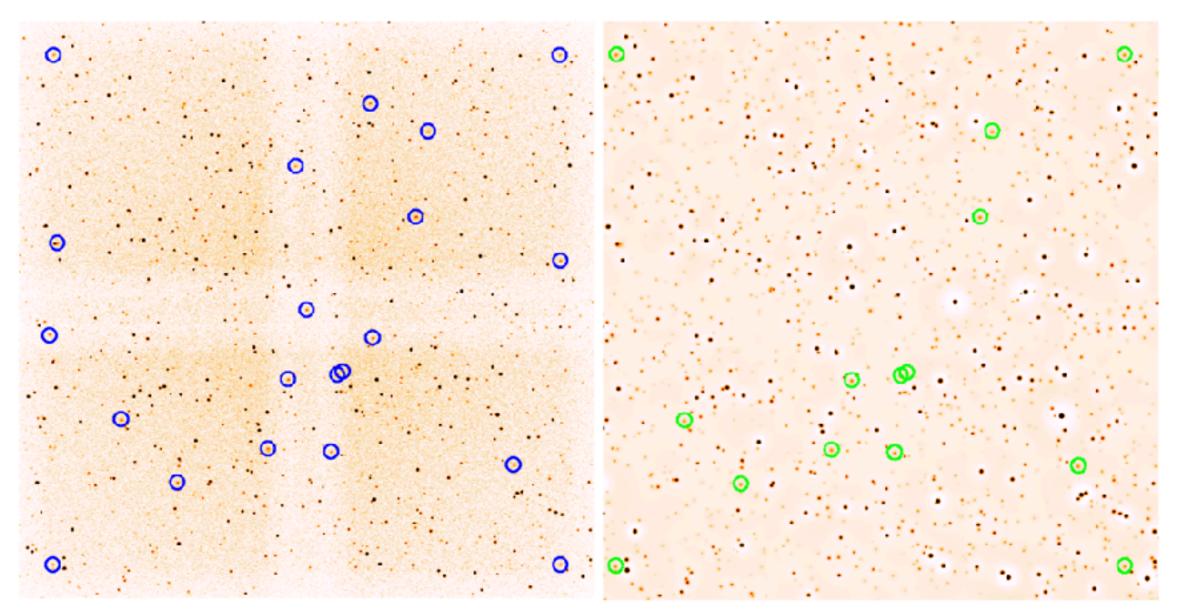


#### Assembly of the first massive halos

z = 2.5 / 80 ks

Simulated

Detected

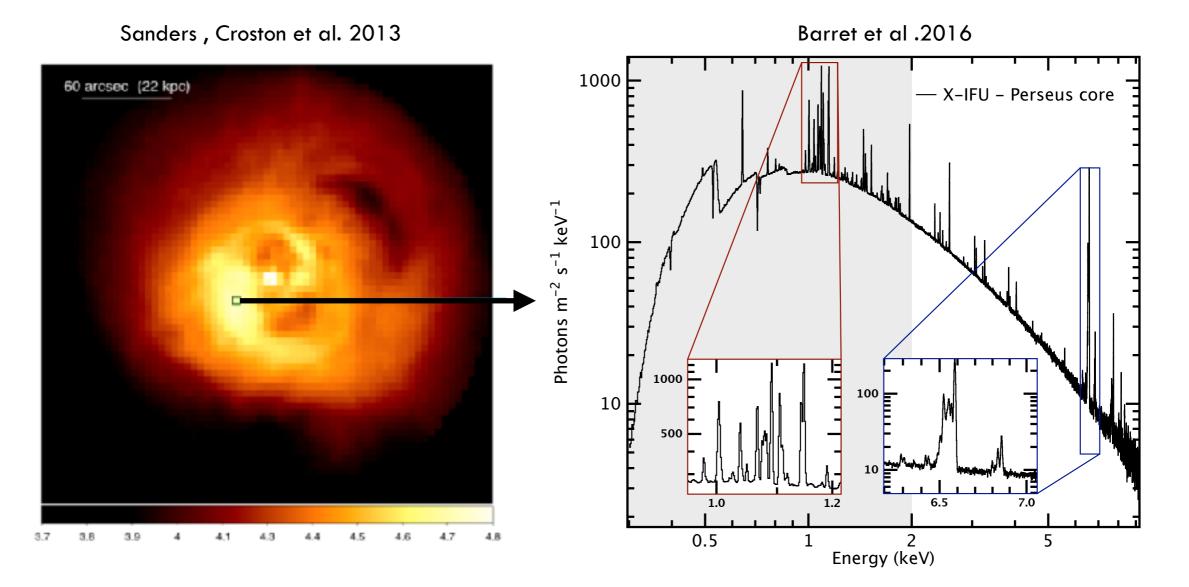


Courtesy of C. Zhang and F. Pacaud



### AGN feedback in clusters

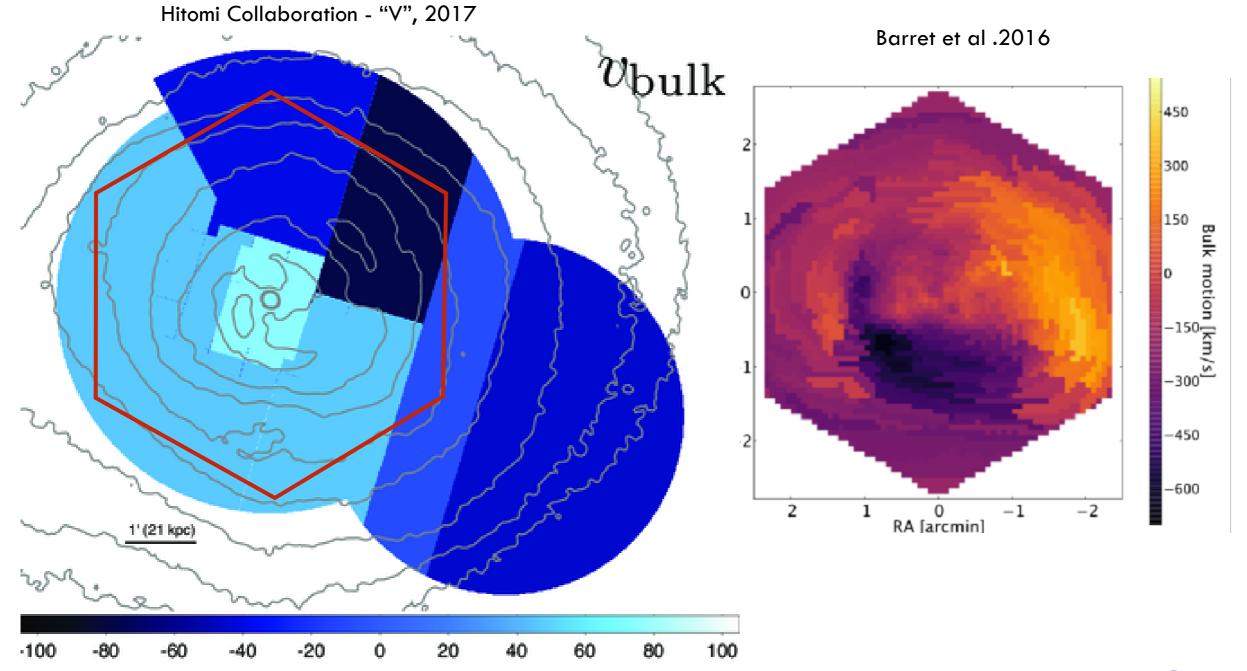
- Energetics of central AGN feedback
- Gas dynamic and jet energy dissipation in clusters





## Dynamical assembly of clusters

- Simulated 'Perseus like' cluster (from Rasia, Biffi. Borgani, Dolag)
- 100ks X-IFU observation with SIXTE



E. Pointecouteau, SnowCluster, 23rd of March 2018



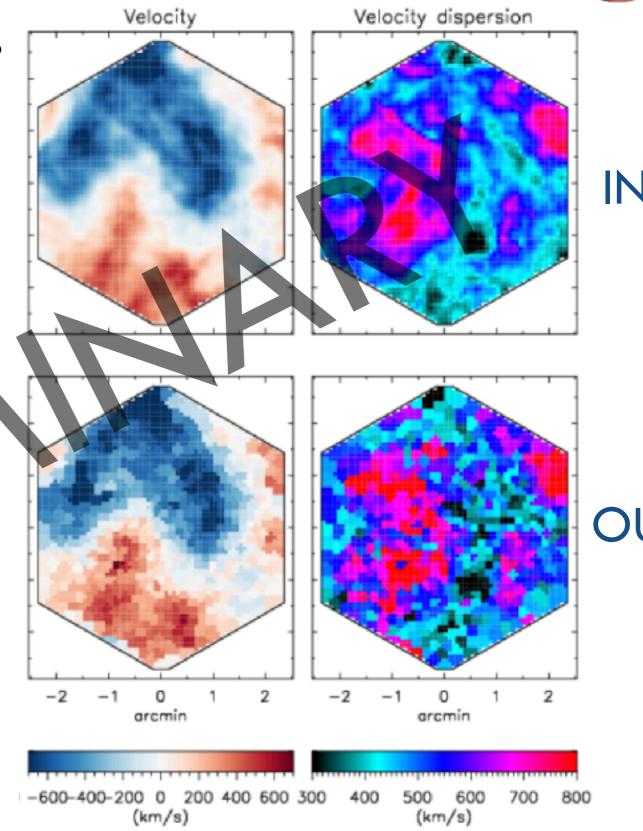
# Dynamical assembly of clusters

Mauro Roncarelli



Roncarelli+ in prep

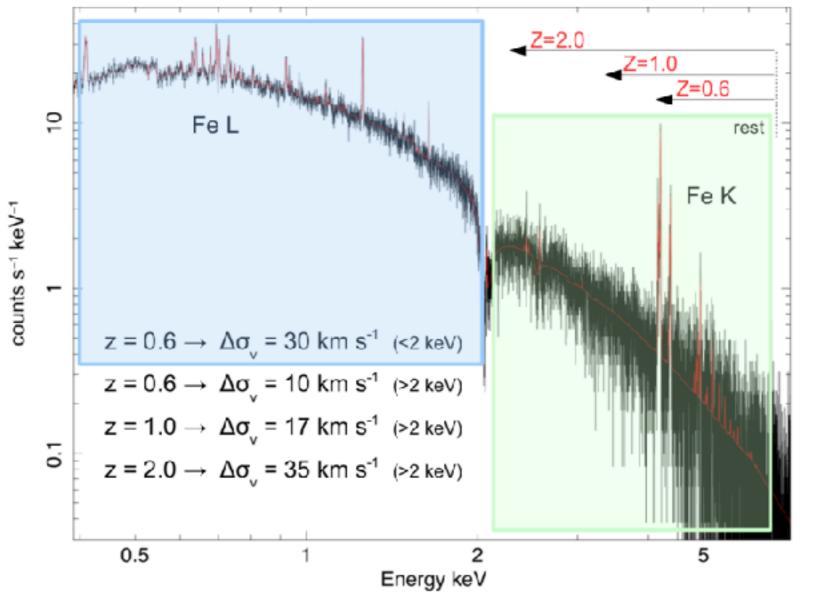
- Simulated Coma like ICM
  - from Gaspari & Churazov 2013
  - turbulence
  - thermal conduction
- X-IFU single pointing
  - SIXTE simulation of observation
  - bapec emission model
- Velocity measurement
  - bulk motion from line shift
  - turbulence from line broadening

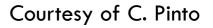




## Dynamical assembly of clusters

- Measuring velocities through line broadening
  - Phoenix cluster at z~0.6
  - ► larcmin diameter area, ~15 ksec, no instrumental noise
  - Statistical errors only





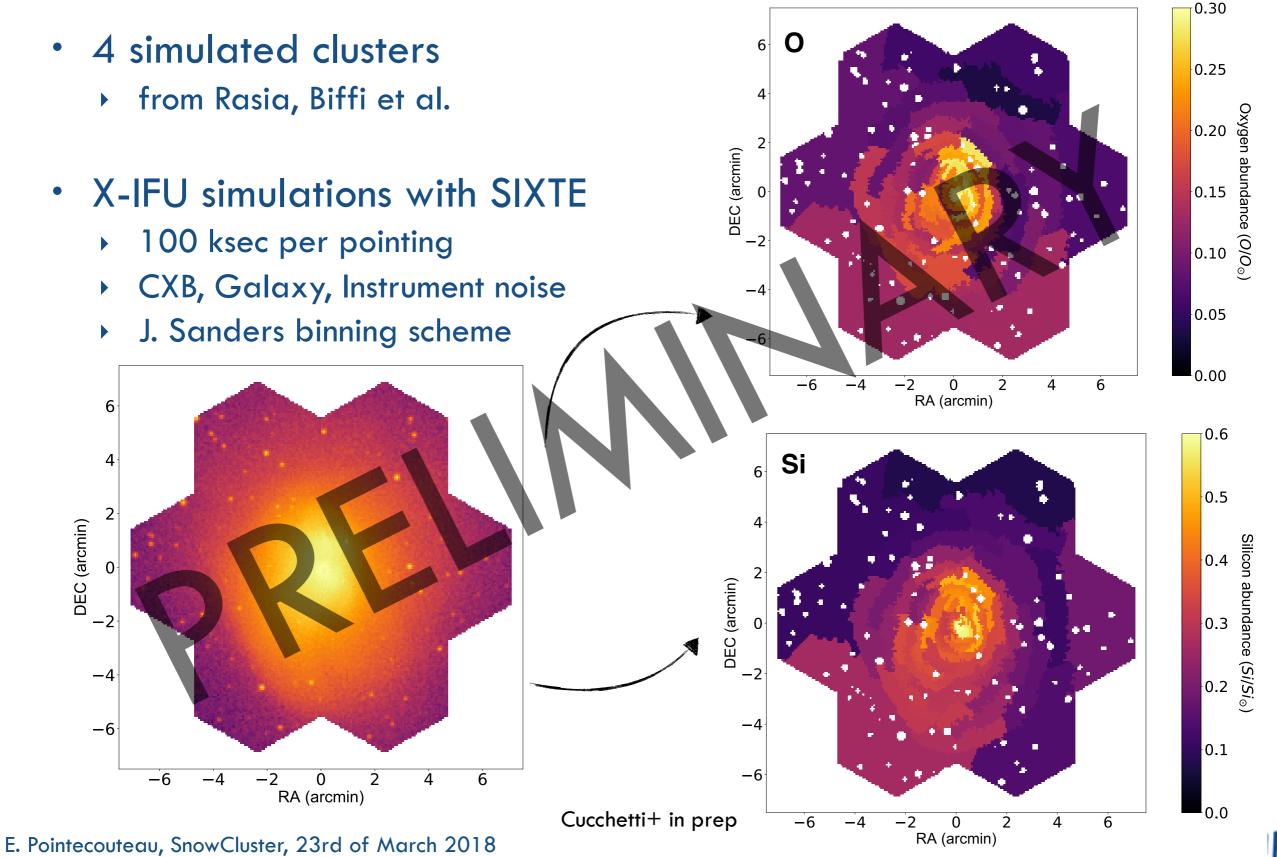




Edoardo Cucchetti

# Chemical evolution





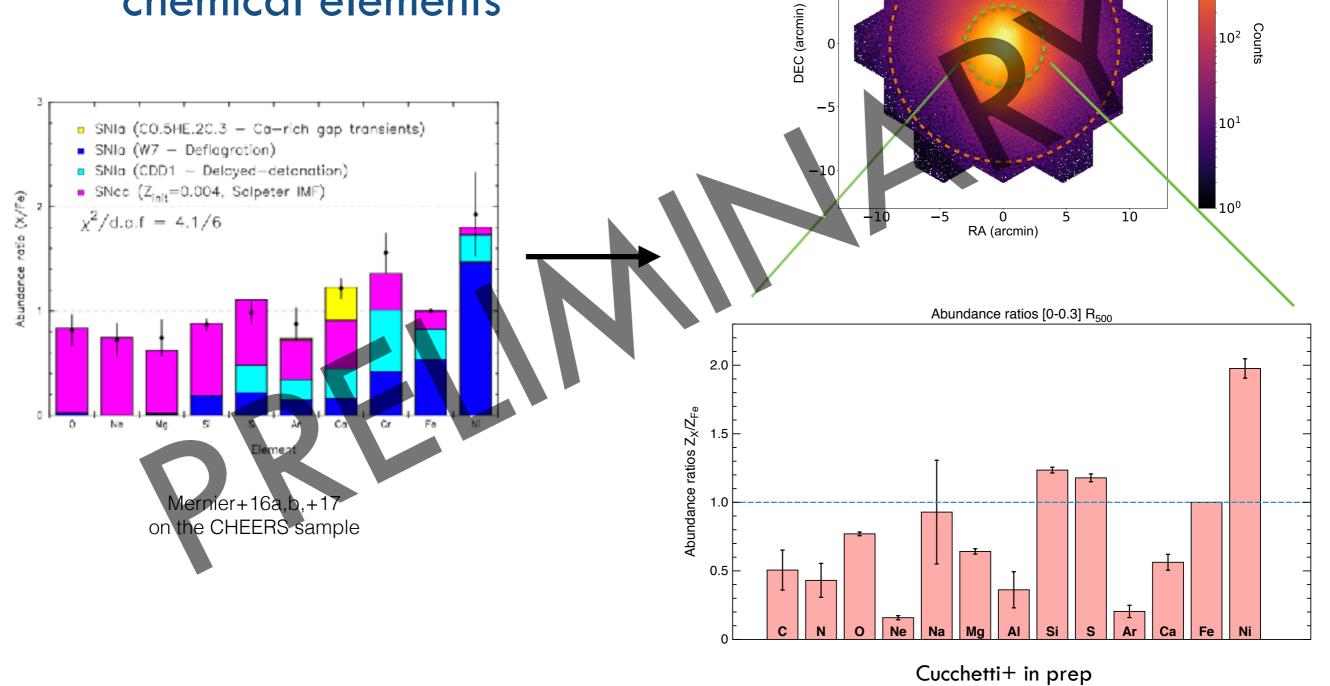
Edoardo Cucchetti



10<sup>3</sup>

# Chemical evolution





10

5

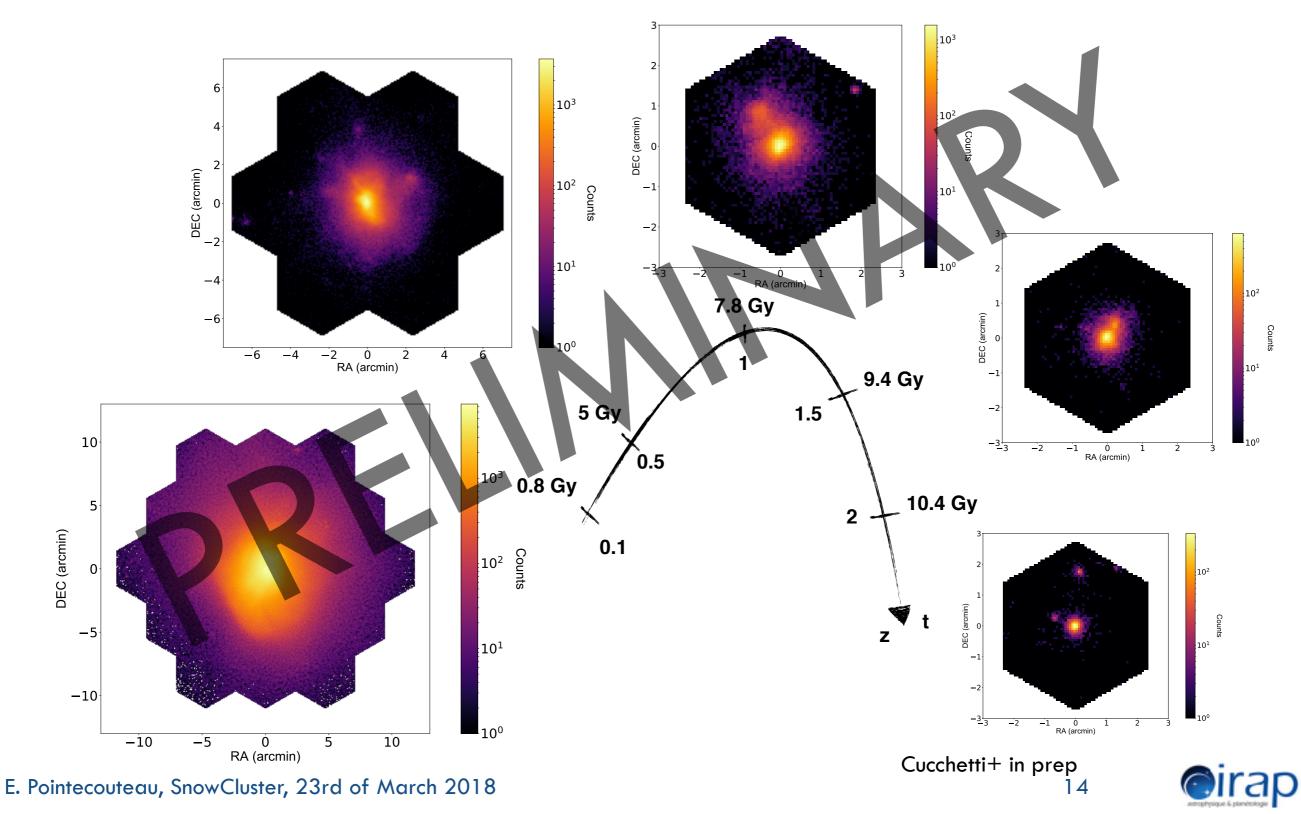


# Chemical evolution

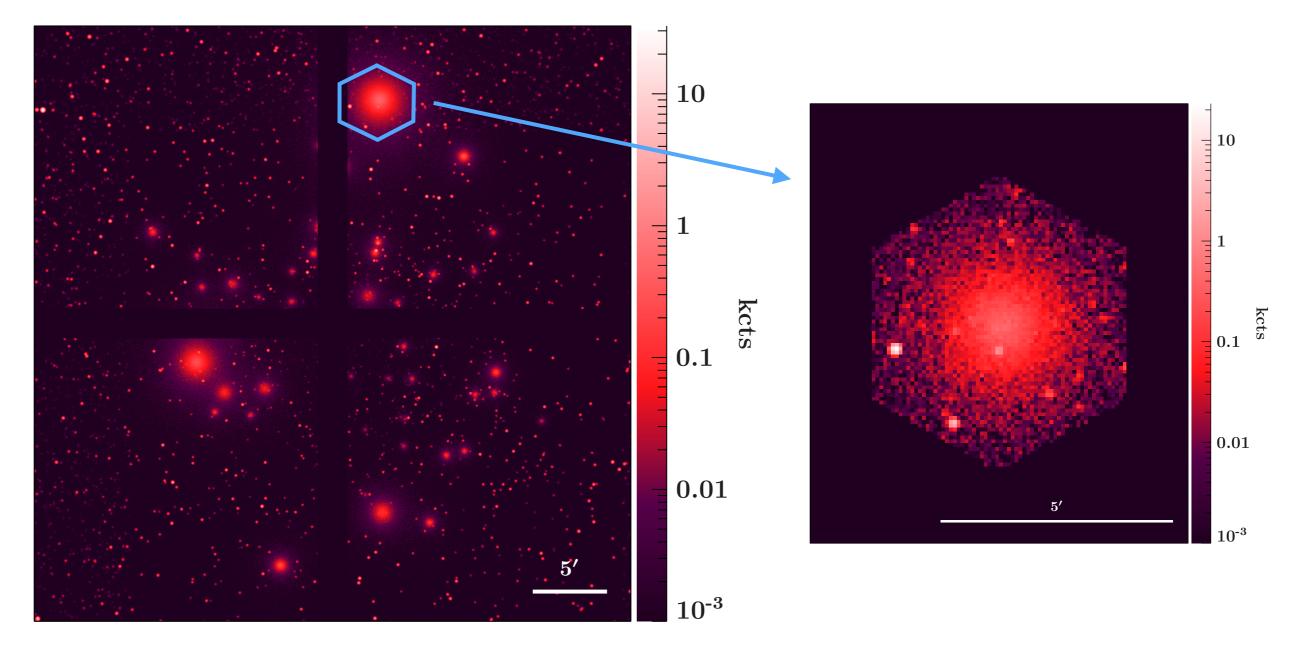
Edoardo Cucchetti



#### • Enrichment of the ICM across cosmic times



#### 150 ksec WFI and X-IFU

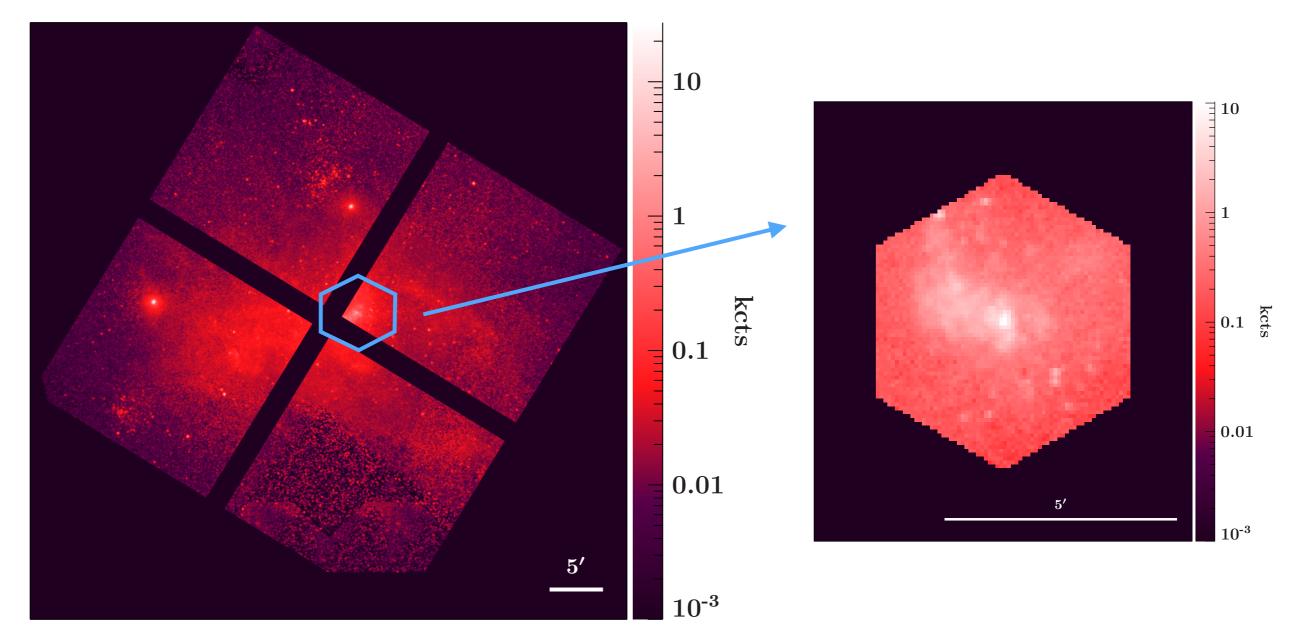


Courtesy of Thomas Dauser and Joern Wilms





#### 150 ksec WFI and X-IFU

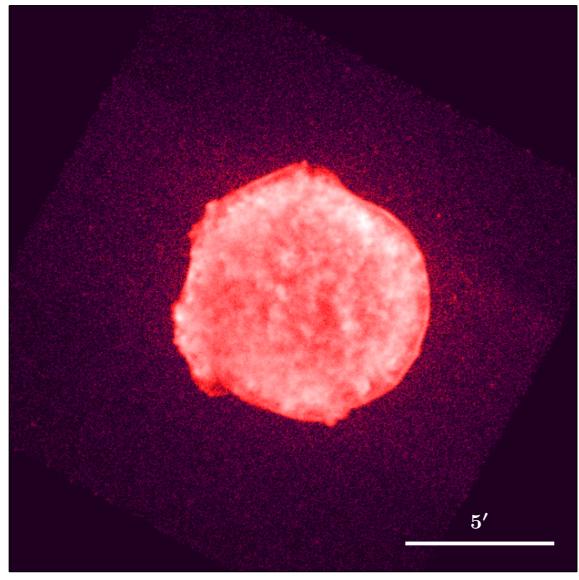


Courtesy of Thomas Dauser and Joern Wilms



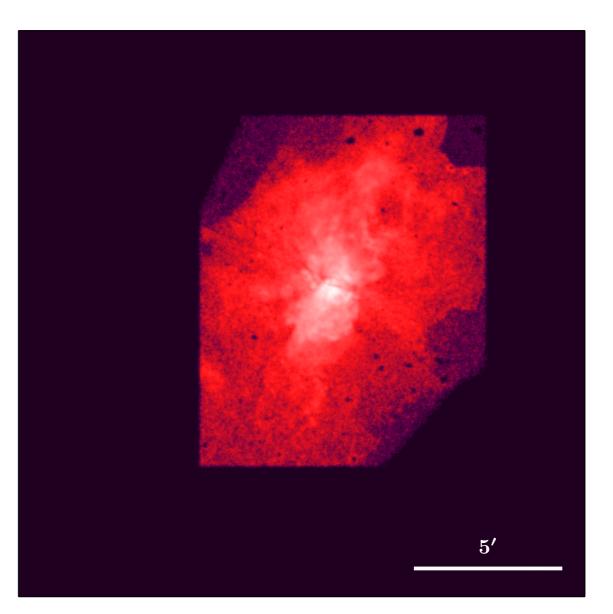
## Galaxies and SNRs

- Tycho
  - WFI 1.5 ksec



Courtesy of Thomas Dauser and Joern Wilms

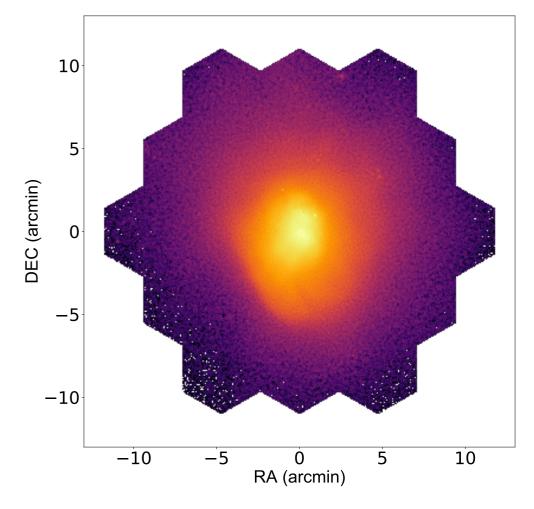
M82
WFI 150 ksec





# Take home messages

- Athena and its instruments are in feasibility/definition phase
- Cluster science is at the core of the Athena science case
  - Driving the telescope and instruments performance (effective area, spatial resolution, energy resolution, energy coverage, FoV, etc)



• Athena will enable a wealth of science on the hot and energetic Universe, and on observatory science



