

ATHENA+

# THE HOT AND ENERGETIC UNIVERSE

1) How does ordinary matter assemble into the large scale structures we see today?

2) How do black holes grow and influence the Universe?"

The Science Theme motivating the Athena+ mission

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Why does the **Observable Universe**  
look the way it does?

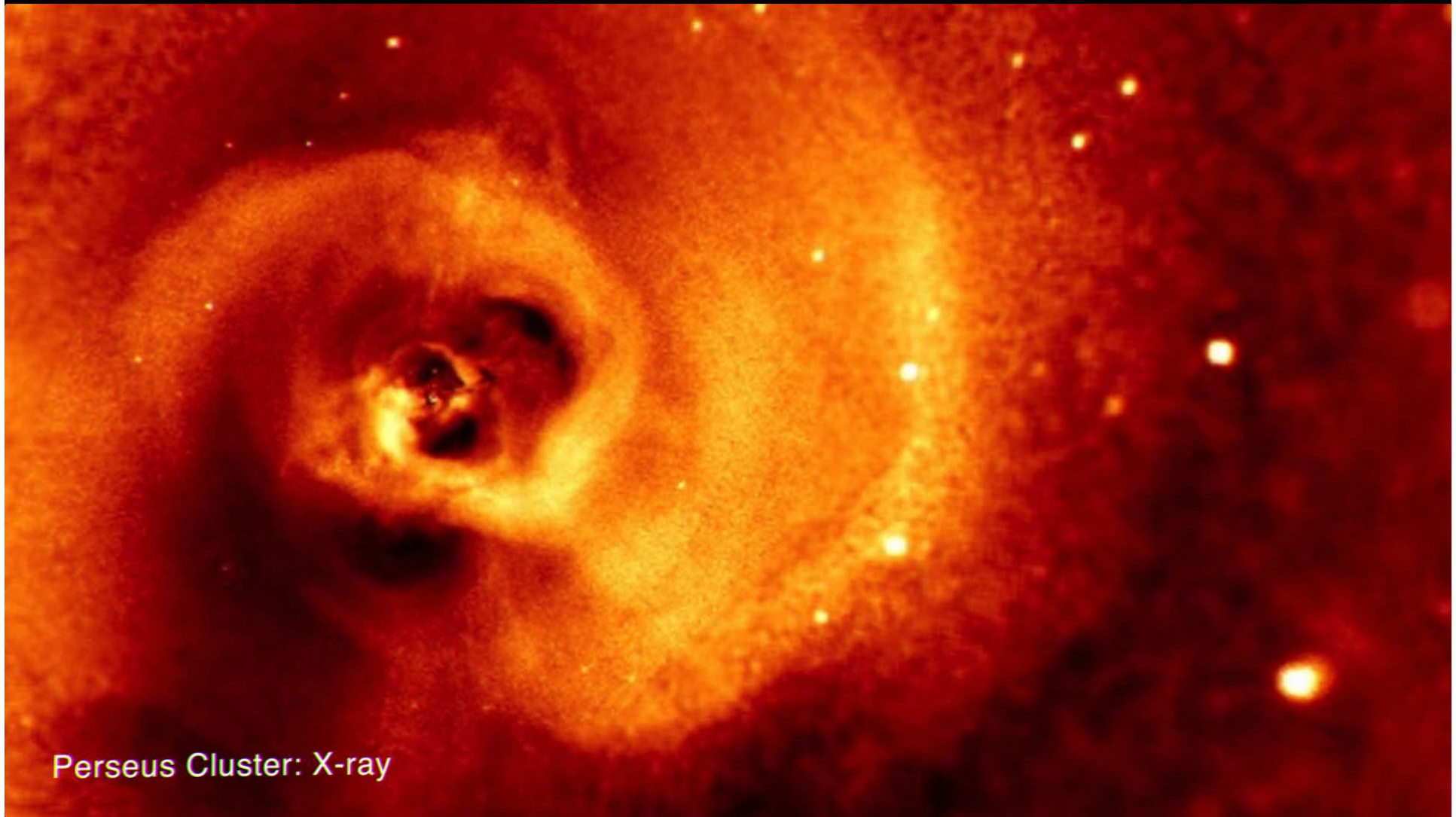
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Why does the **Observable Universe**  
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Perseus Cluster: Optical



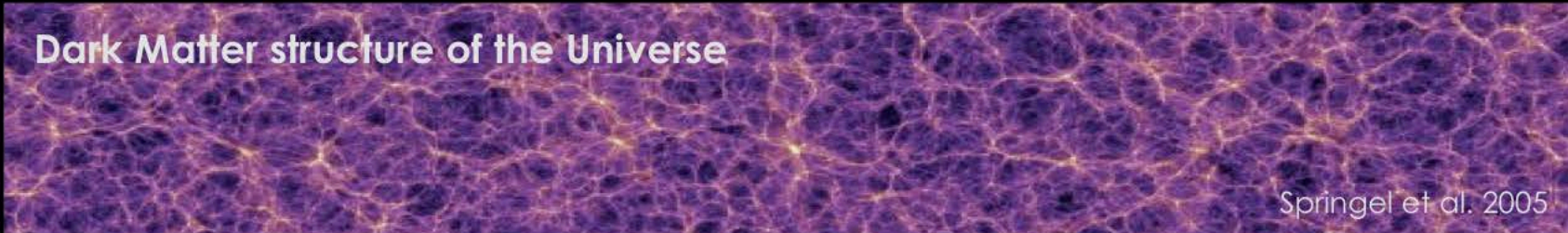
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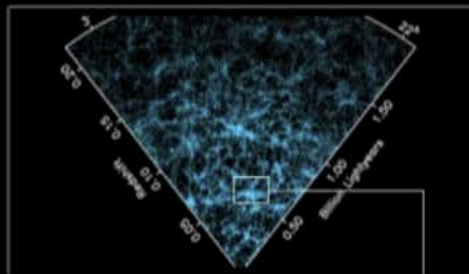
Perseus Cluster: X-ray

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# Why does the observable universe look the way it does?



## THE COLD UNIVERSE

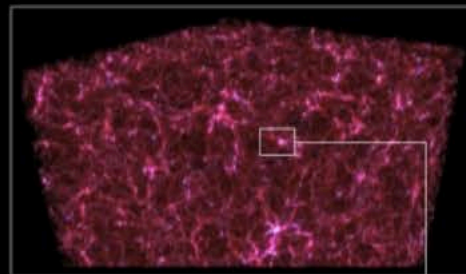


LARGE SCALE GALAXY DISTRIBUTION



STARS AND GALAXIES

## THE HOT UNIVERSE

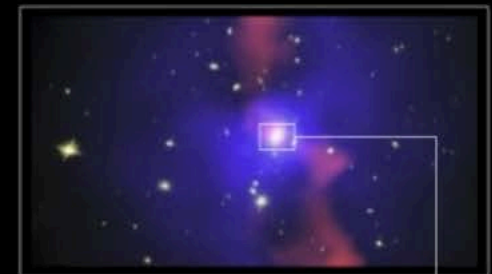


INTERGALACTIC MEDIUM

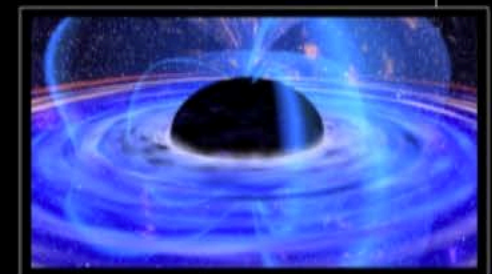


GALAXY CLUSTERS & GROUPS

## THE ENERGETIC UNIVERSE



COSMIC FEEDBACK



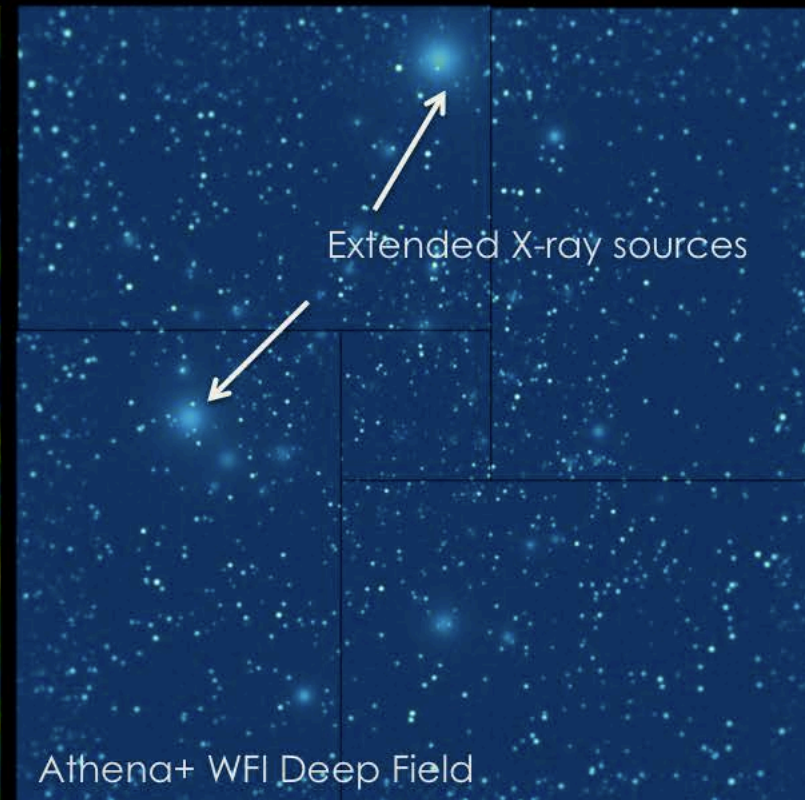
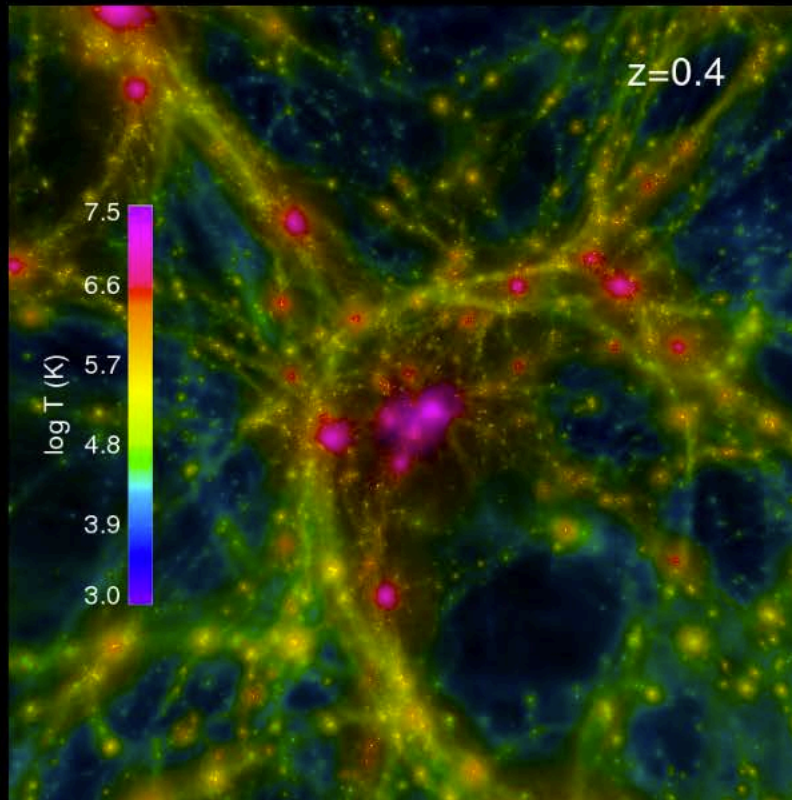
BLACK HOLES



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## Key questions for observational astrophysics in 2028

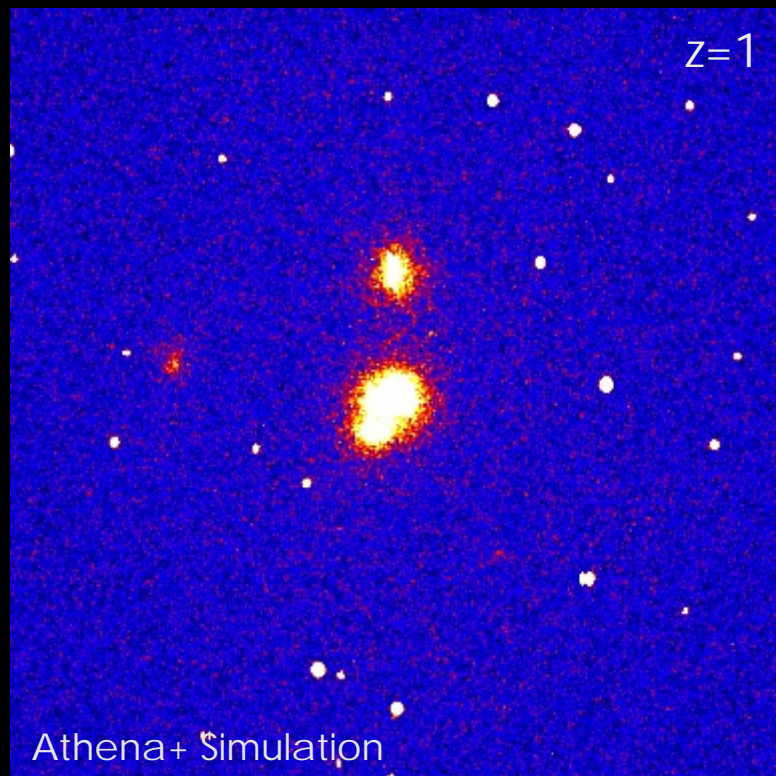
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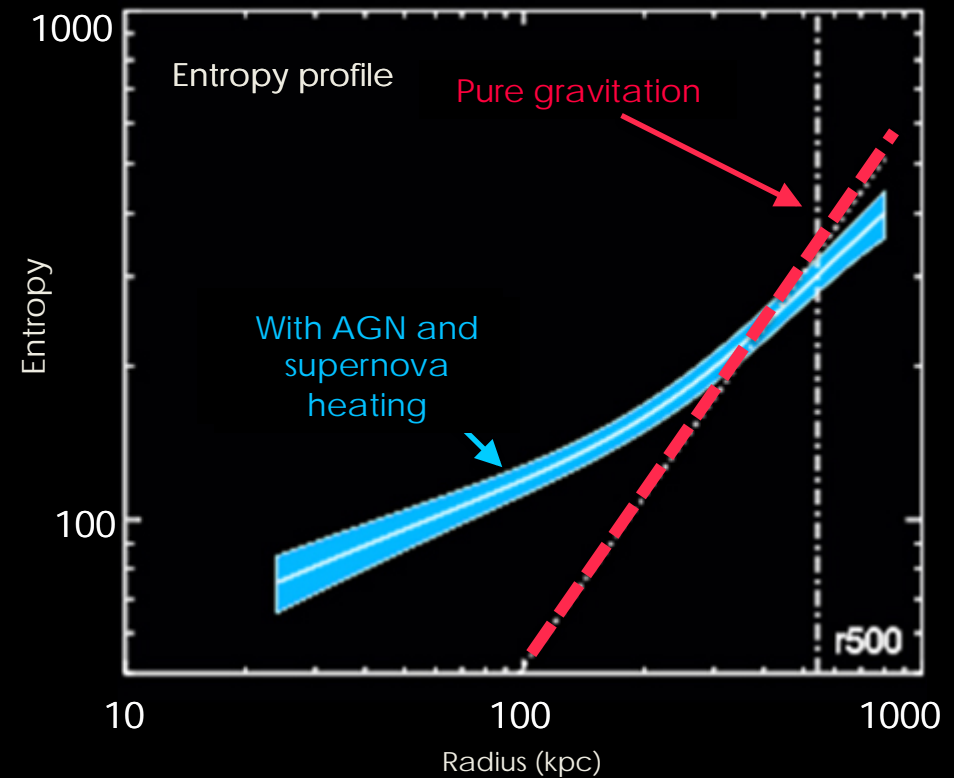
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# The formation and evolution of clusters and groups of galaxies

How and when was the energy contained in the hot intra-cluster medium generated?



Pointecouteau, Reiprich et al., 2013 arXiv1306.2319

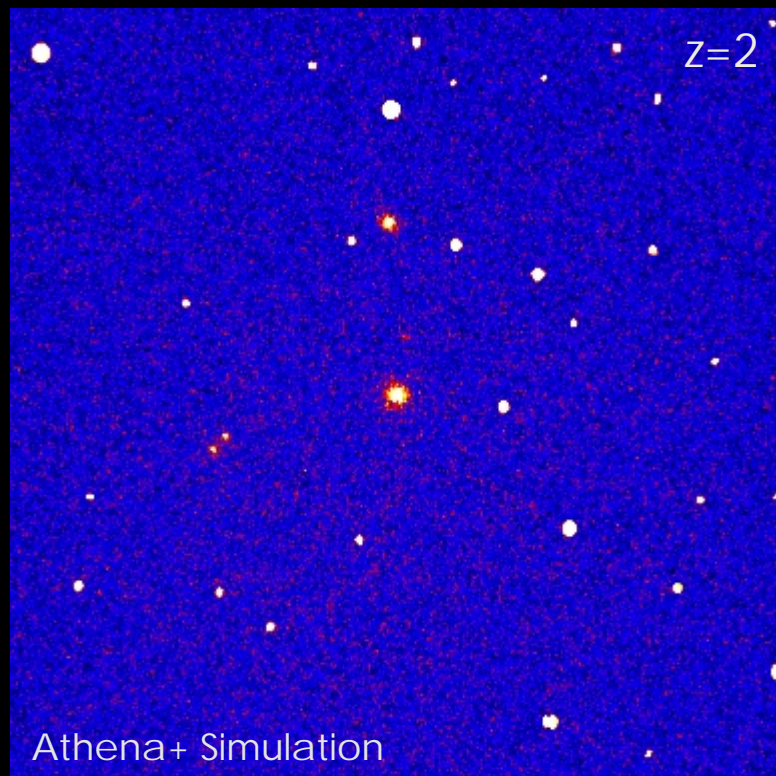


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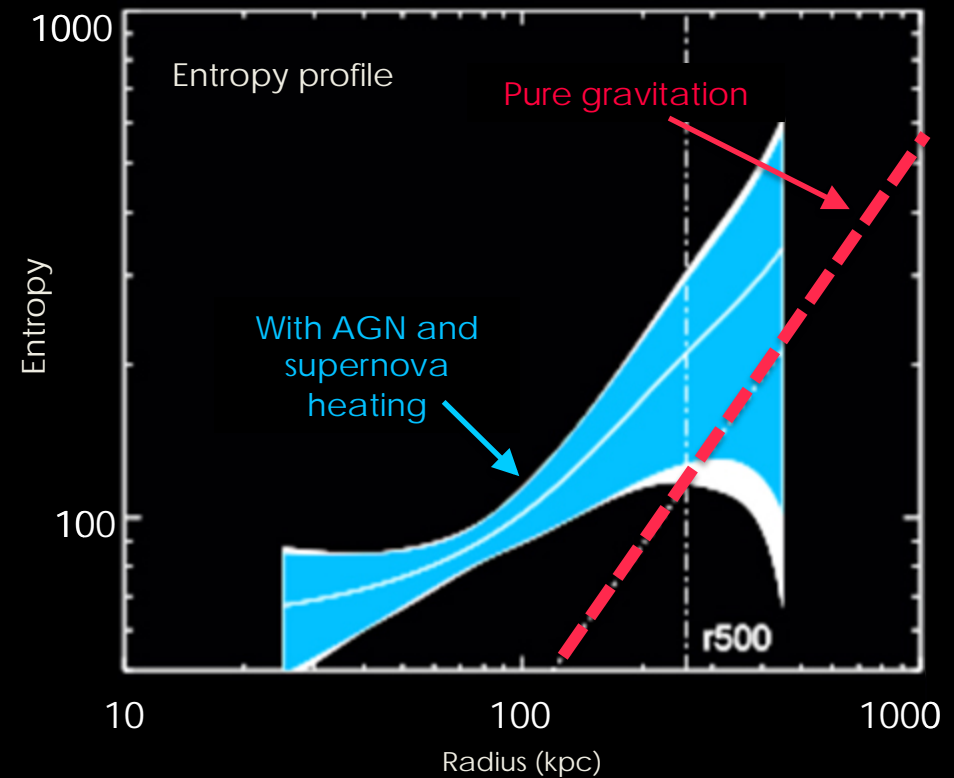
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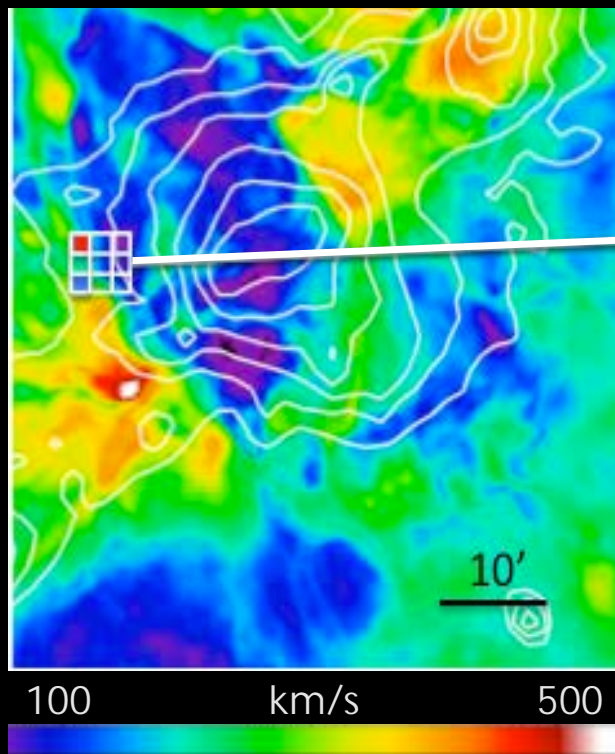


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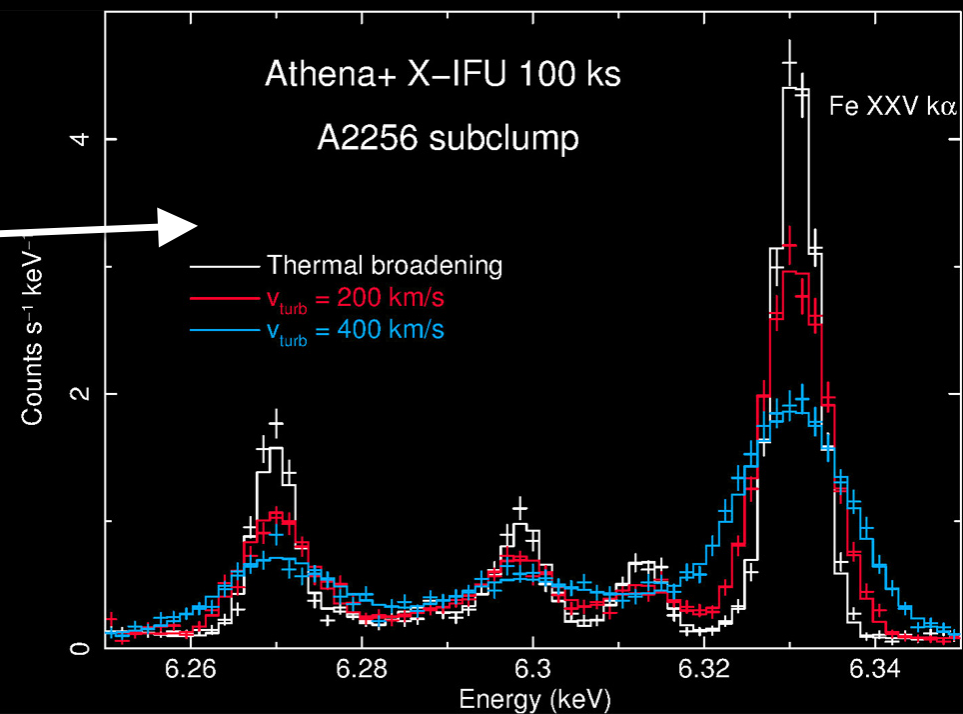
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How and when was the energy contained in the hot intra-cluster medium generated?

Simulated Velocity map



Ettori, Pratt, et al., 2013 arXiv1306.2322

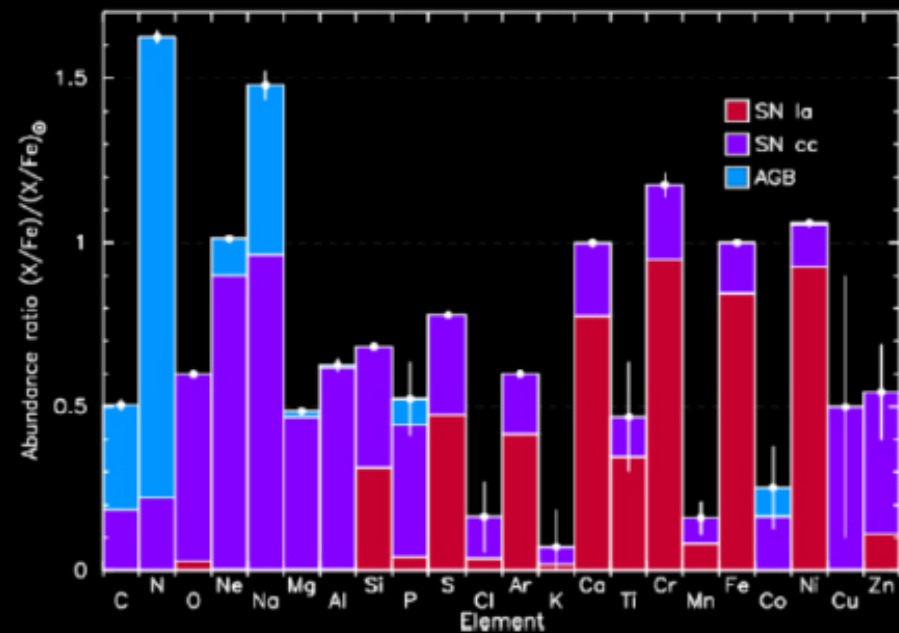
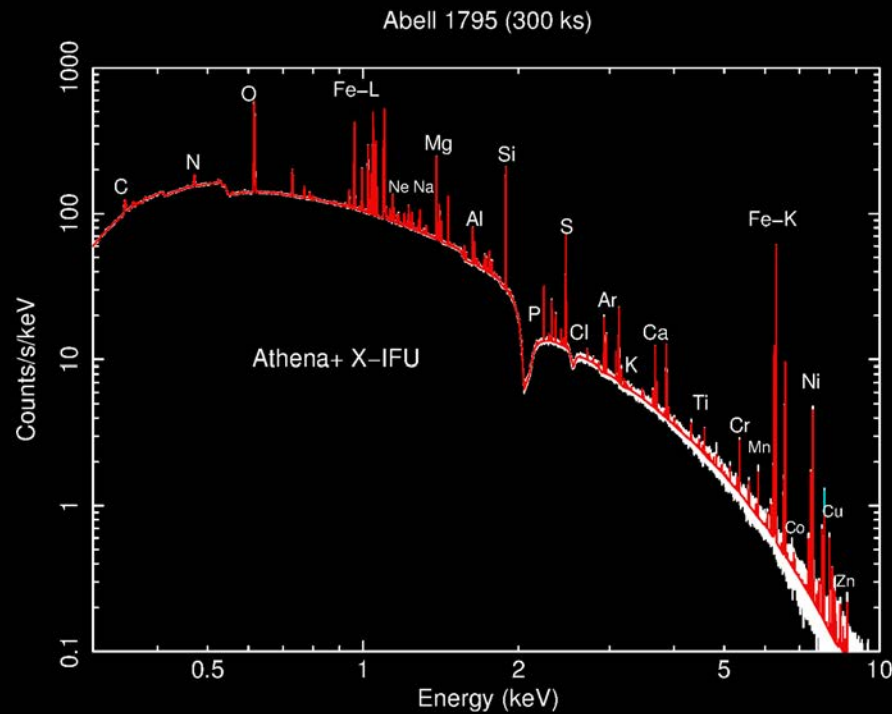


How does ordinary matter assemble into the large-scale structures that we see today?

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# The chemical evolution of hot baryons

When and how were the largest baryon reservoirs in galaxy clusters chemically enriched?



Ettori, Pratt, et al., 2013 arXiv1306.2322

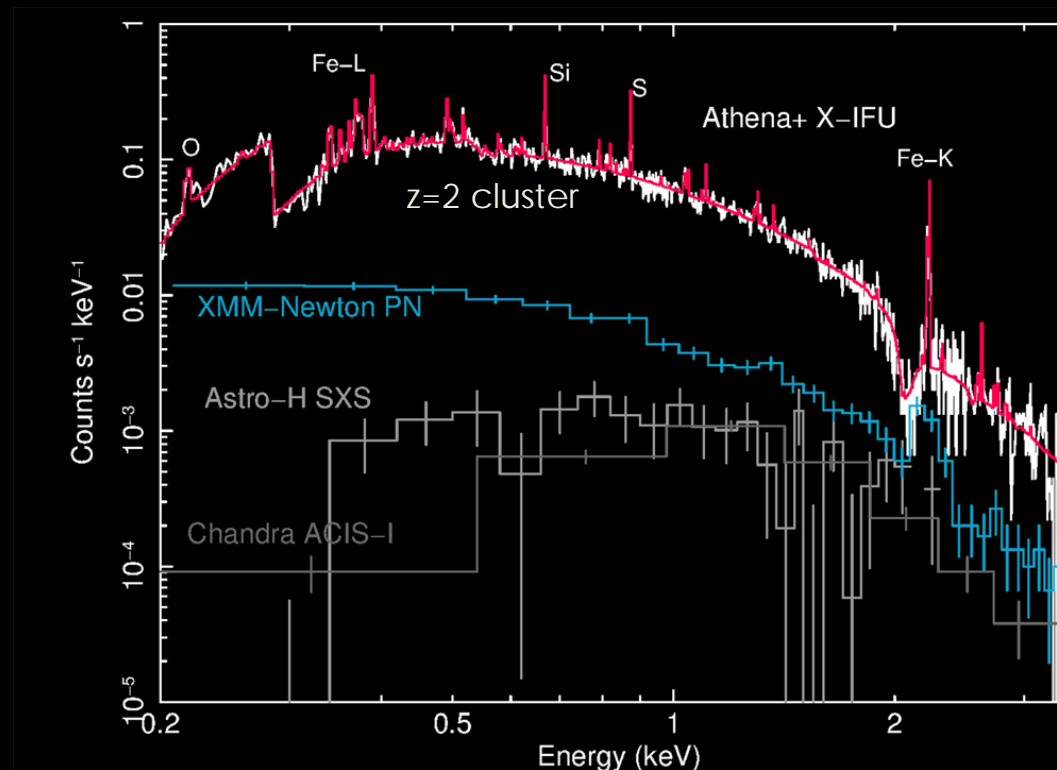
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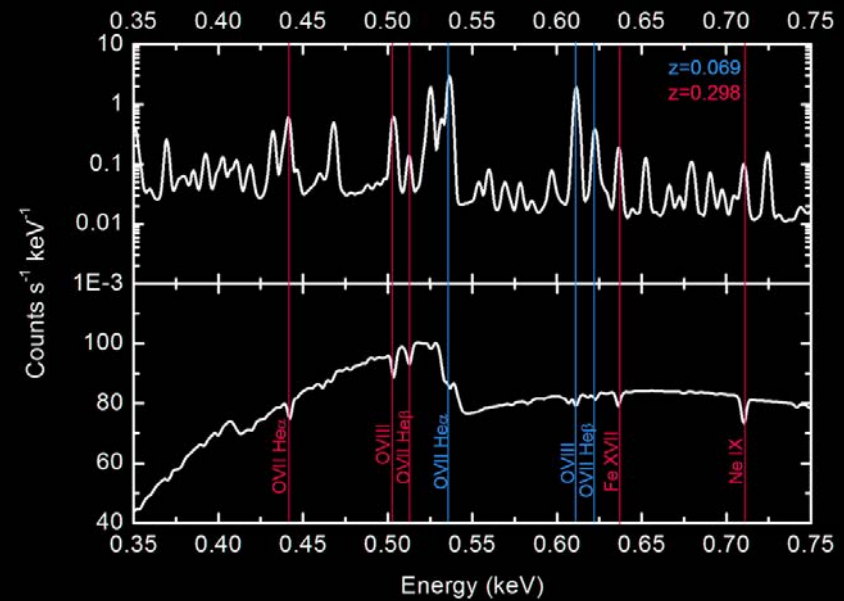
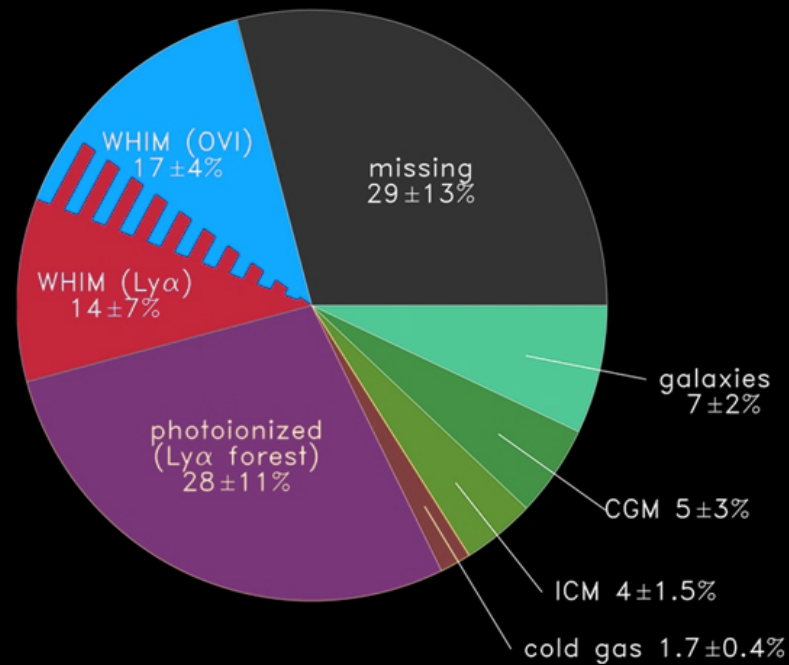
Ettori, Pratt, et al., 2013 arXiv1306.2322

How does ordinary matter assemble into the large-scale structures that we see today?

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# The Warm-Hot intergalactic medium (WHIM)

Where are the missing baryons in the local Universe? What is the underlying mechanism determining the distribution of the hot phase of the cosmic web?



Kaastra, Finoguenov et al., 2013 arXiv1306.2324

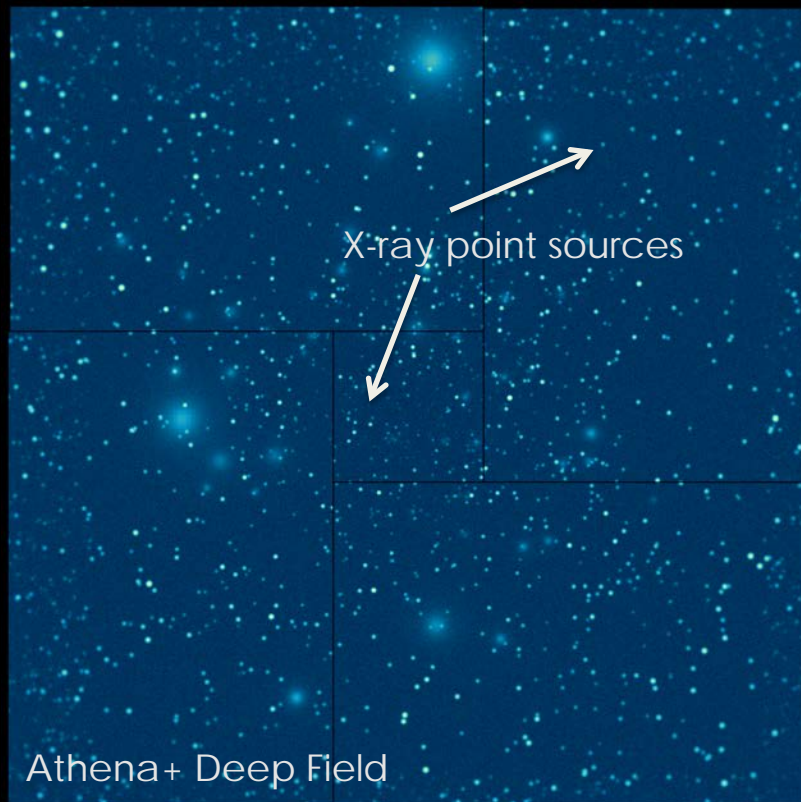
How does ordinary matter assemble into the large-scale structures that we see today?



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## Key questions for observational astrophysics in 2028

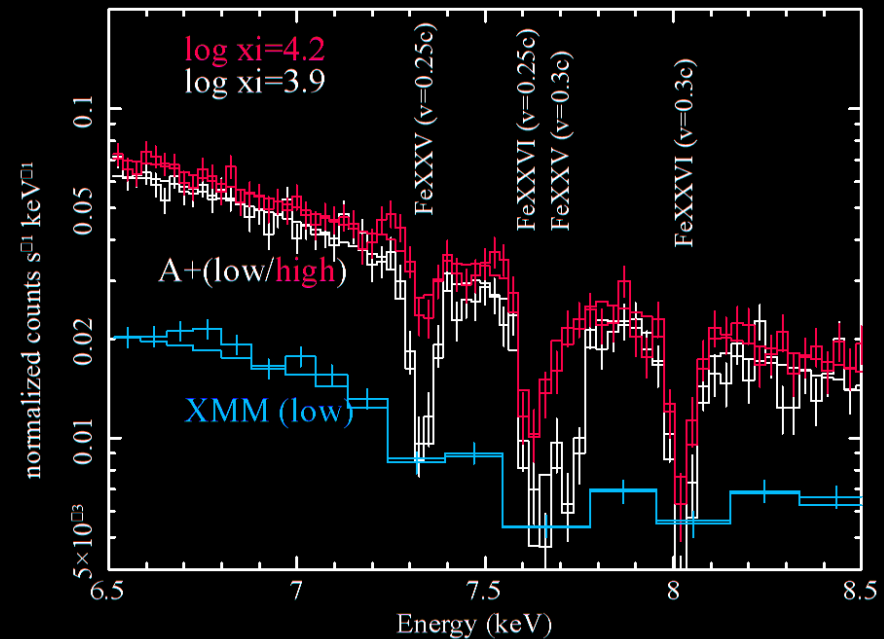
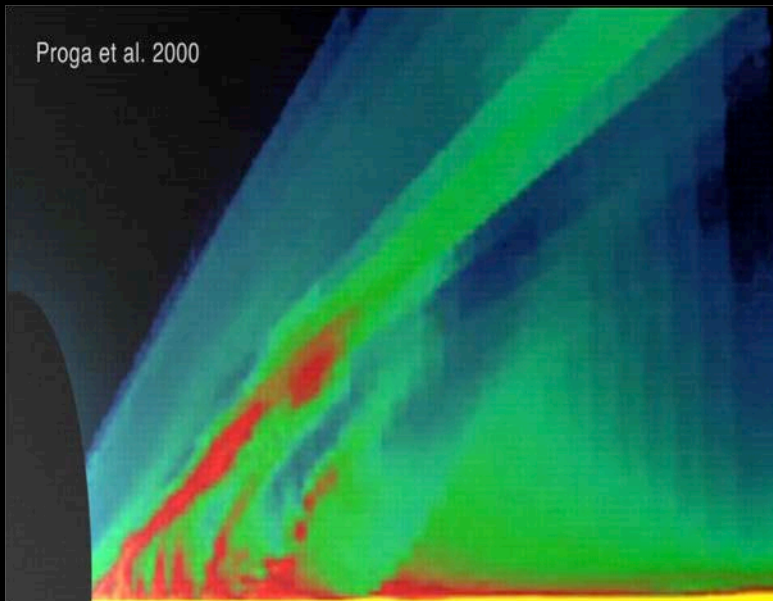
1. How does ordinary matter assemble into the large scale structures we see today?
2. How do black holes grow and shape the Universe?



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# Cosmic feedback: the origin of black hole winds

How do black holes launch winds and outflows?  
How much energy do they carry out to larger scales?



Cappi, Done et al., 2013 arXiv1306.2330  
Dovciak, Matt et al., 2013 arXiv1306.2331

How do black holes grow and shape the Universe?

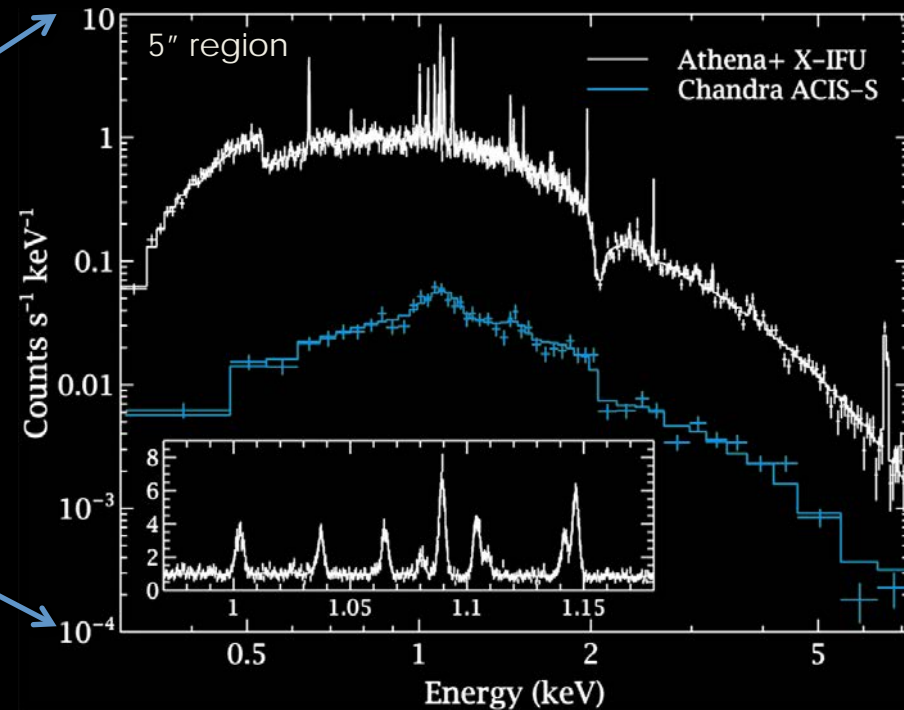
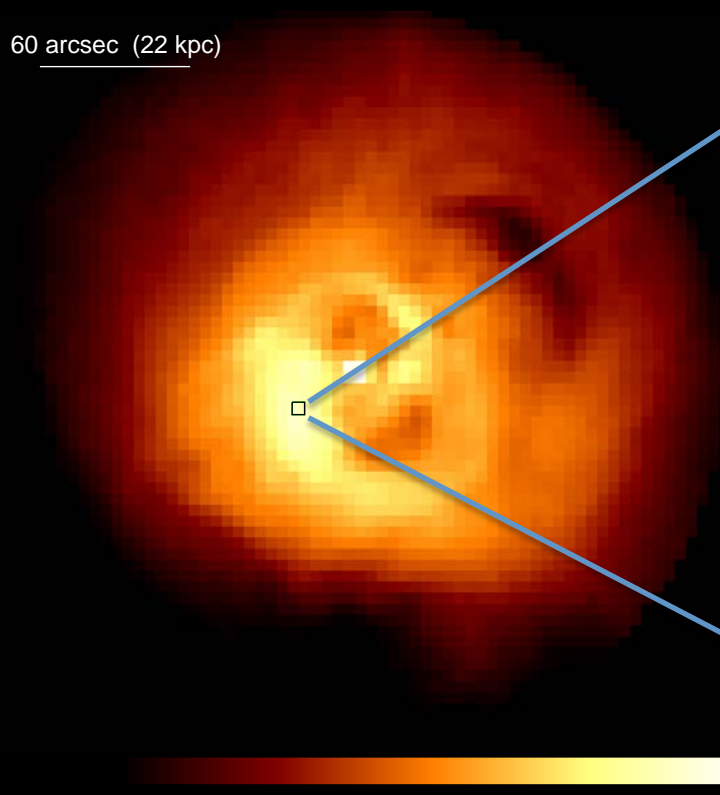


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# Cosmic feedback: the impact on galaxy cluster scales

How do jets from Active Galactic Nuclei dissipate their mechanical energy in the hot intracluster medium, and how does this regulate gas cooling and black hole fuelling?

60 arcsec (22 kpc)



Croston, Sanders et al., 2013 arXiv1306.2323

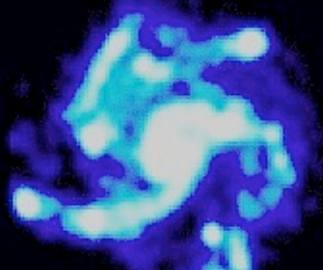
How do black holes grow and shape the Universe?

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# Cosmic feedback: black hole and galaxy co-evolution

How much black hole accretion occurs in the most obscured environments?  
How does this relate to the evolution of the host galaxy?

Disk instability

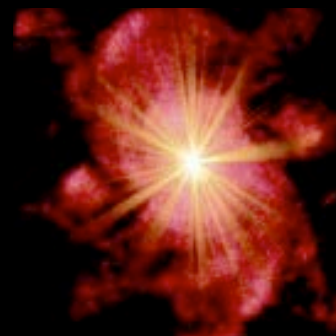


Ceverino et al. 2010

Obscured BH growth



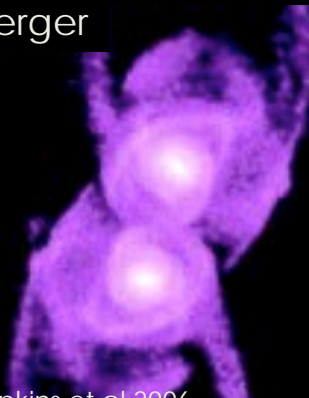
Feedback phase



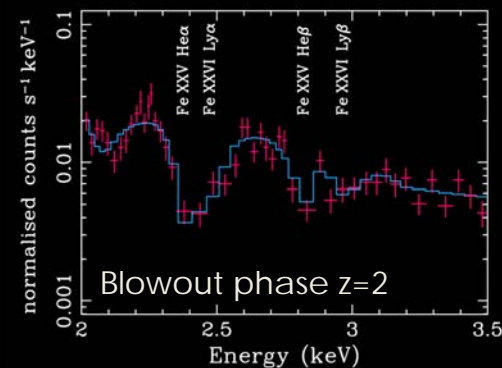
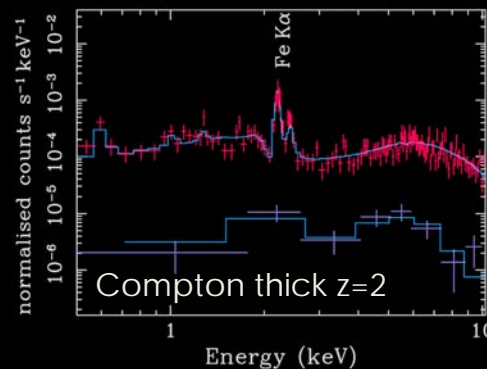
Quiescent remnant



Merger



Hopkins et al. 2006



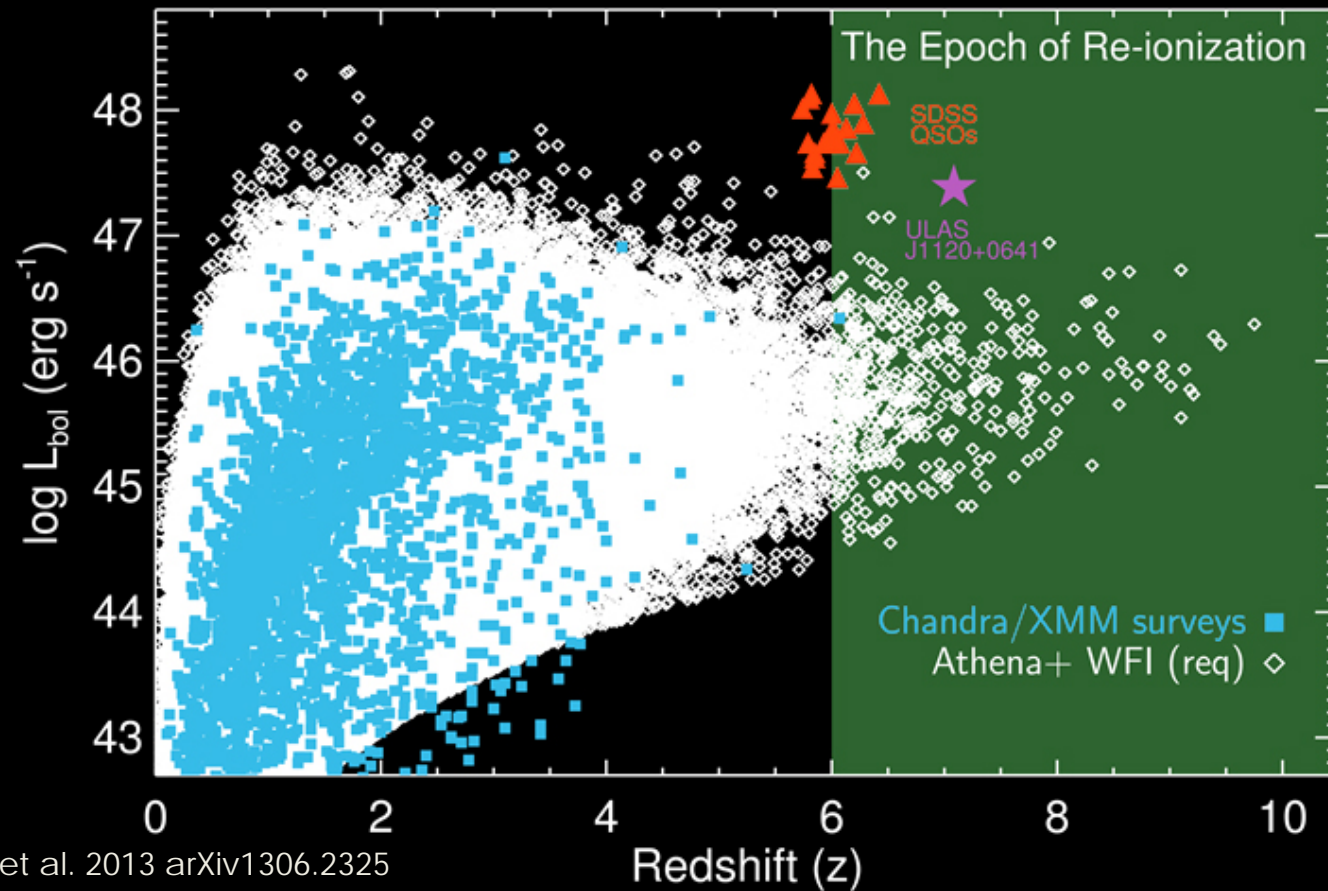
Georgakakis, Carrera et al., 2013 arXiv1306.2328

How do black holes grow and shape the Universe?

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# Black hole growth in the early Universe

What was the growth history of black holes in the epoch of reionization?

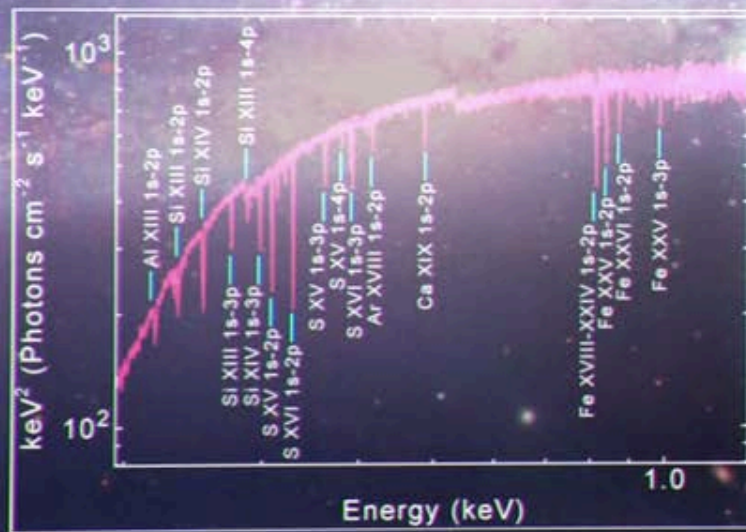


Aird, Comastri et al. 2013 arXiv1306.2325

How do black holes grow and shape the Universe?



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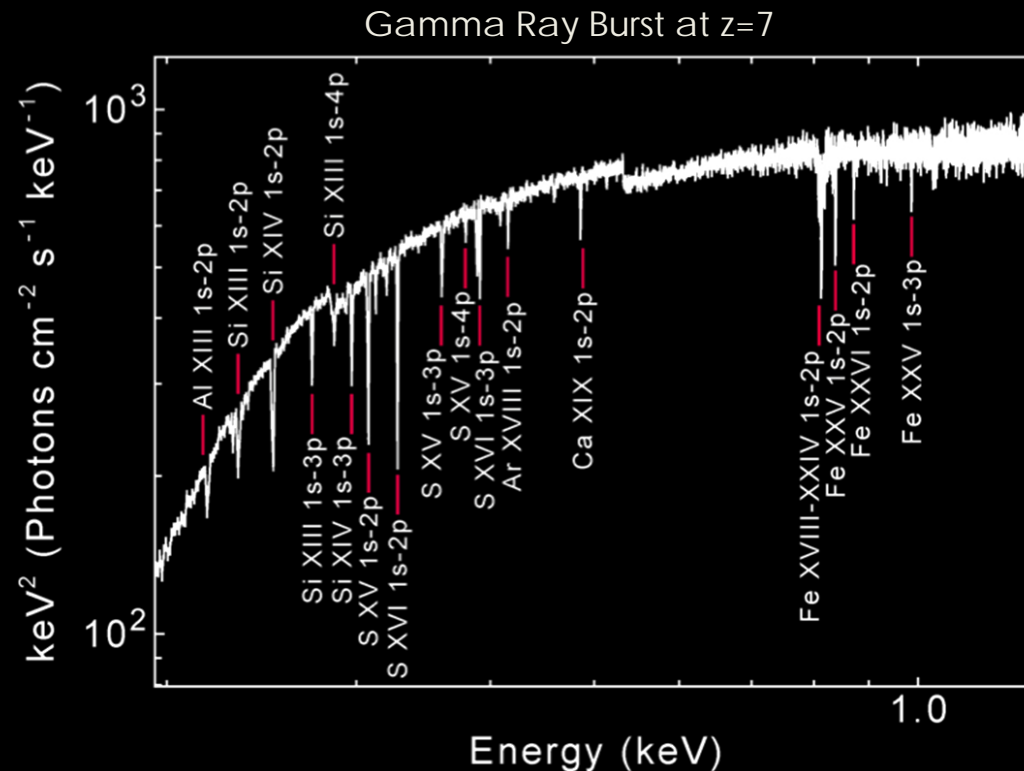


Gamma-ray burst at  $z=7$



# The first stars and black holes

When did the first generation of stars explode to form the first seed black holes and disseminate the first metals in the Universe?

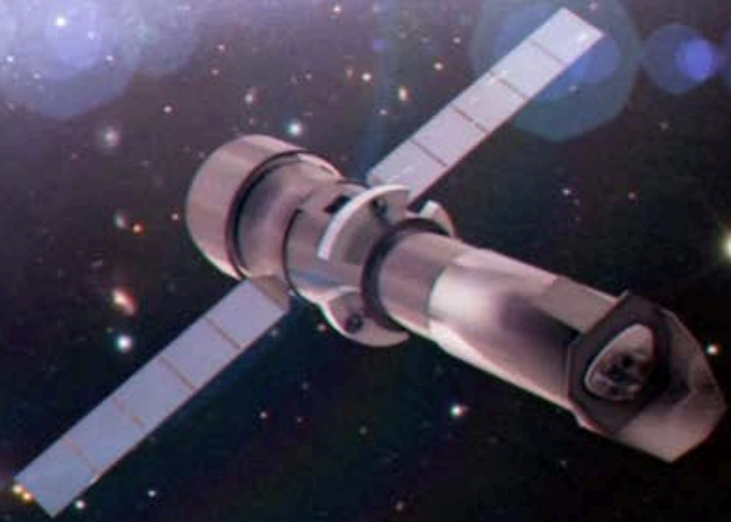


Jonker, O'Brien et al., 2013 arXiv1306.2336

How do black holes grow and shape the Universe?

ATHENA+

THE ATHENA+  
OBSERVATORY



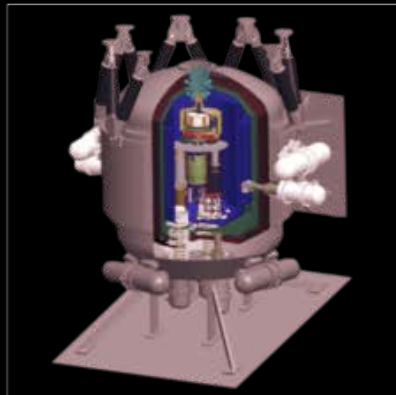


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# The Athena+ Observatory

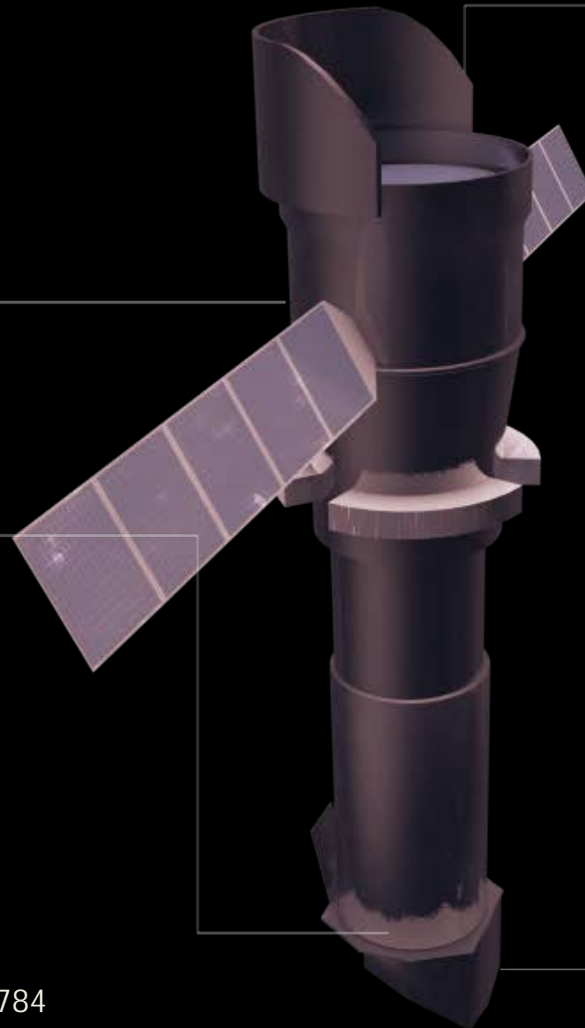
Willingale et al, 2013  
arXiv1308.6785

L2 orbit Ariane V  
Mass < 5100 kg  
Power 2500 W  
5 year mission

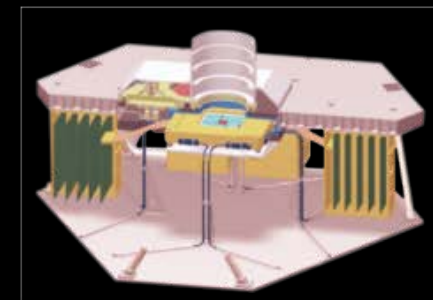


X-ray Integral Field Unit:  
 $\Delta E$ : 2.5 eV  
Field of View: 5 arcmin  
Operating temp: 50 mk

Barret et al., 2013 arXiv:1308.6784



Silicon Pore Optics:  
2 m<sup>2</sup> at 1 keV  
5 arcsec HEW  
Focal length: 12 m  
Sensitivity:  $3 \cdot 10^{-17}$  erg cm<sup>-2</sup> s<sup>-1</sup>



Wide Field Imager:  
 $\Delta E$ : 125 eV  
Field of View: 40 arcmin  
High countrate capability

Rau et al. 2013 arXiv1307.1709

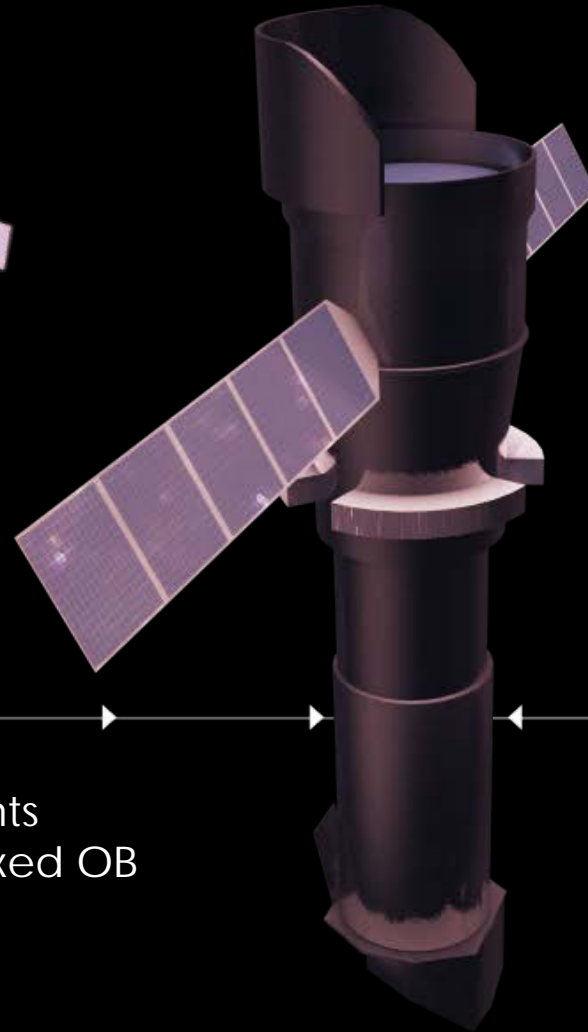
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## Technical Maturity

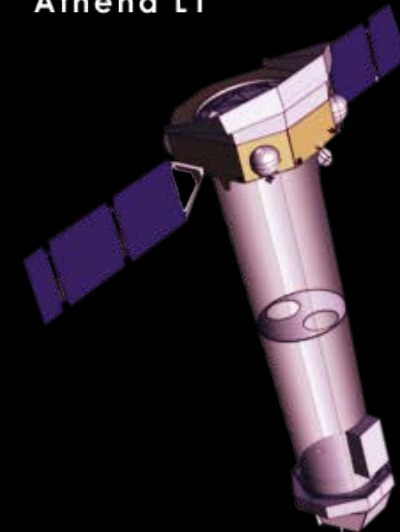
IXO (Ariane 5)



Simplified:  
5 to 2 instruments  
Extendible to fixed OB



Athena L1

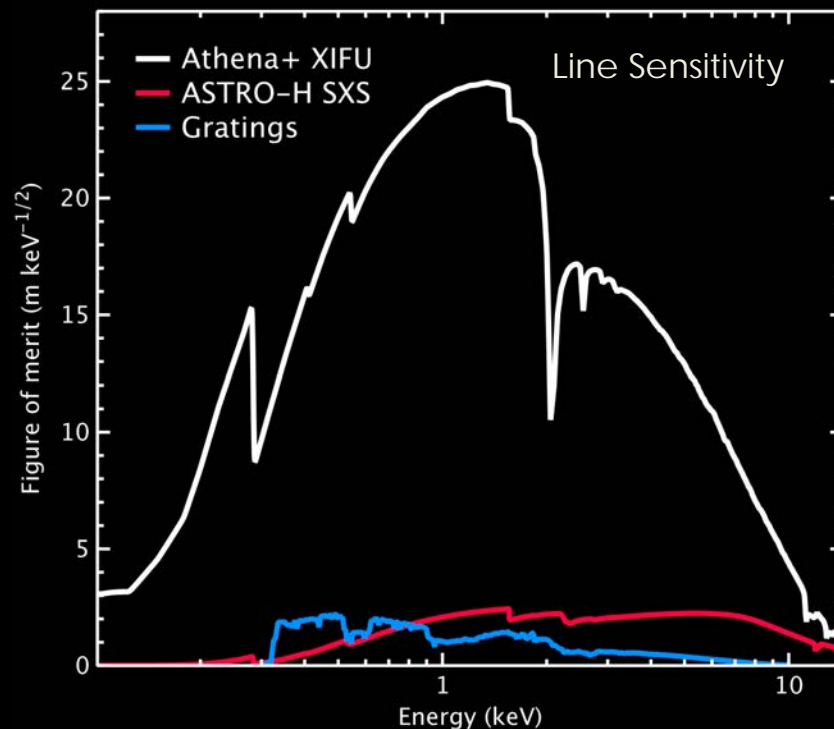


Enhanced:  
Angular resolution now 5"  
Fields of view increase x 4  
Effective area increase x 4  
(per instrument)

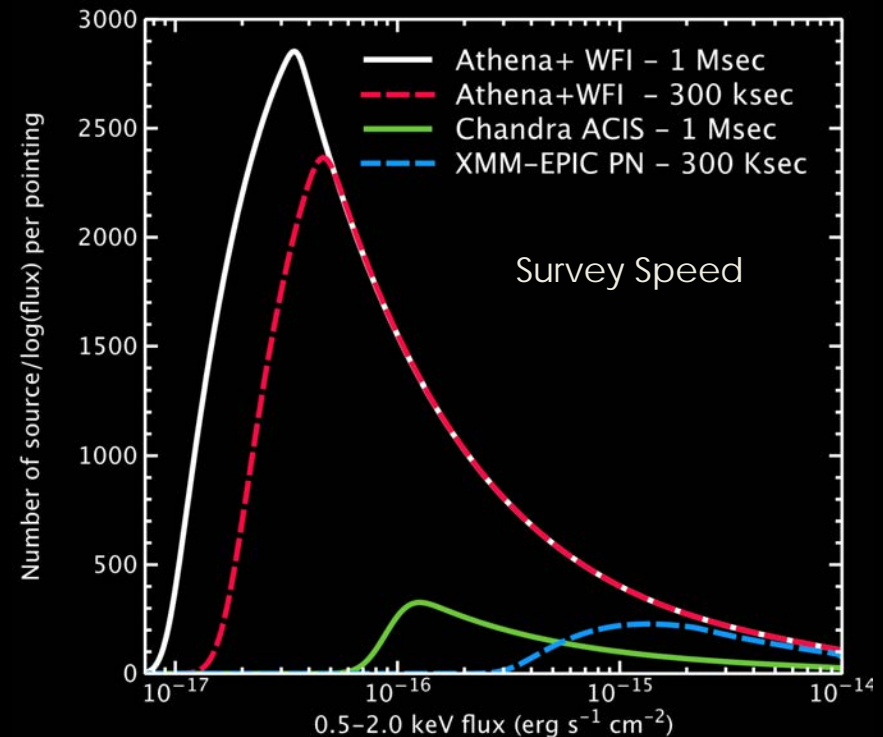
ATHENA+

# The first Deep Universe X-ray Observatory

Athena+ has vastly improved capabilities compared to current or planned facilities, and will provide transformational science in virtually all areas of astrophysics



X-ray spectroscopy at the peak of the activity of the Universe



Deep survey capability into the dark ages and epoch of reionization



ATHENA+

# Athena+: a powerful observatory

Planets

(interaction of solar wind with planet environment and comets)

Exoplanets

Stellar physics

Supernovae

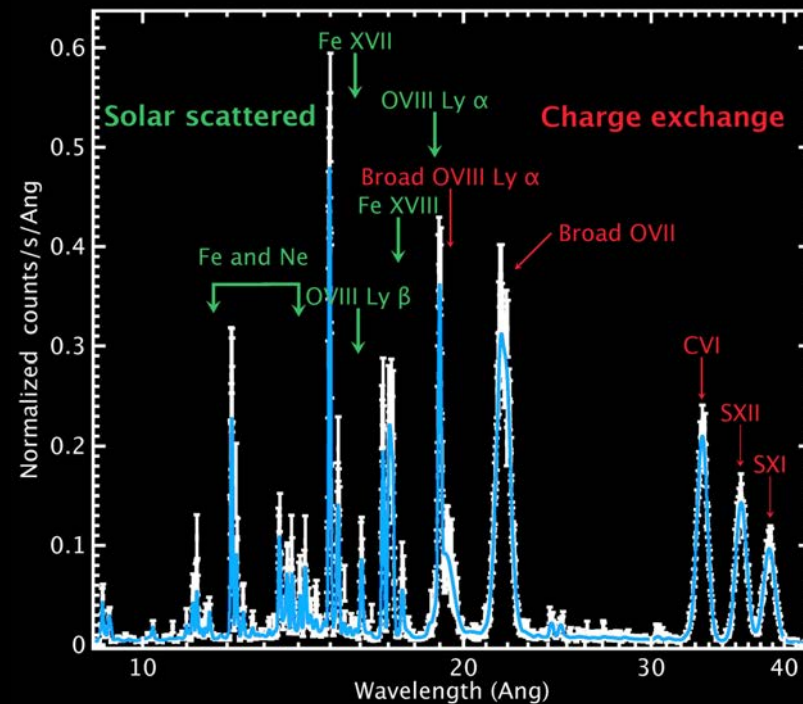
(explosion mechanism, heavy element production)

Stellar endpoints

(physics of outflows and winds in X-ray binaries)

Sgr A\*

Interstellar dust and medium



Jupiter

Branduardi-Raymont, Sciortino, et al., 2013 arXiv 1306.2332; Sciortino, Rauw et al., 2013 arXiv1306.2333; Motch, Wilms, et al., 2013 arXiv1306.2334; Decourchelle, Costantini et al., 2013 arXiv1306.2335

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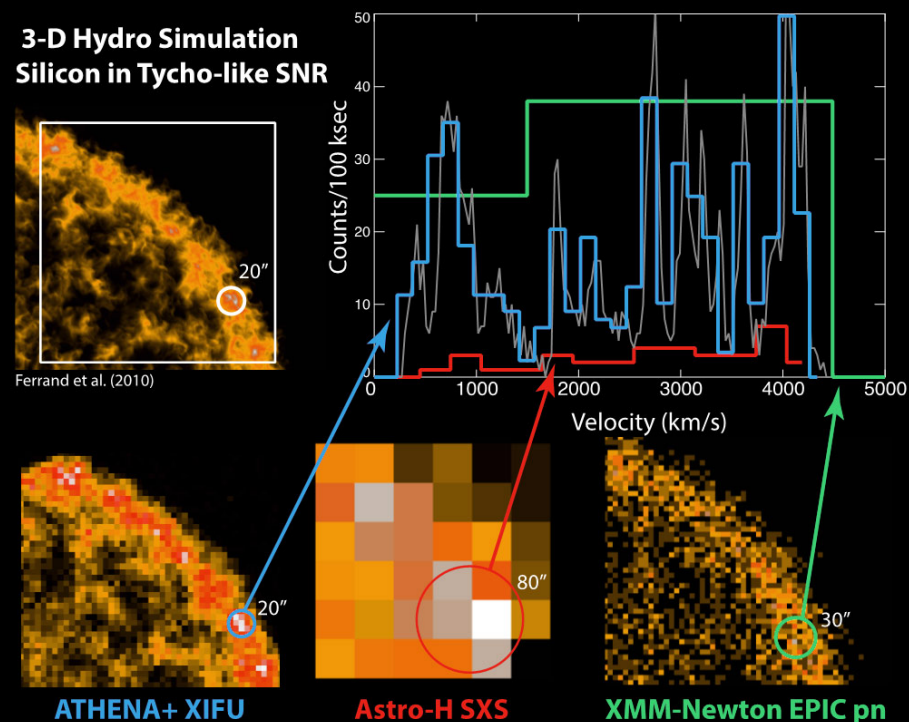
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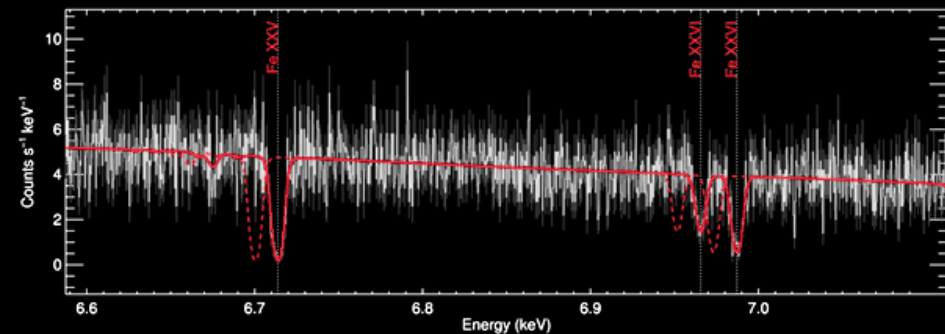
Stellar endpoints

(physics of outflows and winds in X-ray binaries)

Sgr A\*

Interstellar dust and medium

Outflow in X-ray binary, 10ks

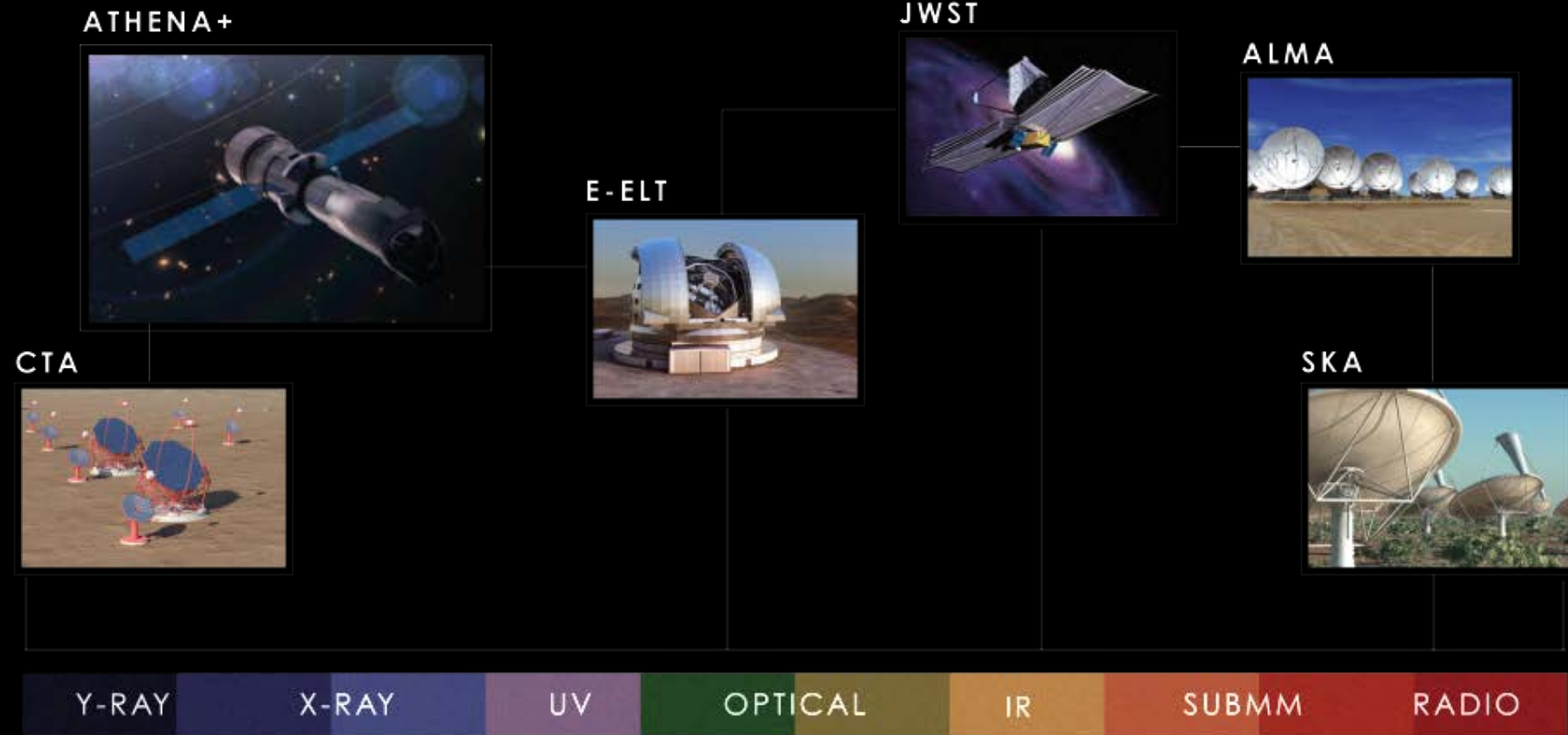


Branduardi-Raymont, Sciortino, et al., 2013 arXiv 1306.2332; Sciortino, Rauw et al., 2013 arXiv1306.2333;  
Motch, Wilms, et al., 2013 arXiv1306.2334; Decourchelle, Costantini et al., 2013 arXiv1306.2335



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# Athena+ science in context



Athena+ is a crucial part of the suite of large observatories needed to reach the science objectives of astronomy in the coming decades

# Acknowledgements

The Athena+ Co-ordination Group: Xavier Barcons (ES), Didier Barret (FR), Andy Fabian (UK), Jan-Willem den Herder (NL), Kirpal Nandra (DE), Luigi Piro (IT), Mike Watson (UK)

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Bold Face Denotes Working Group Chairs

**Athena+ Coordination Group:**  
**K. Nandra, D. Barret, X. Barcons, J.-W. den Herder, A. Fabian, L. Piro, M. Watson**

**Athena+ Working Groups**  
 (~250 people)

**Athena+ supporters**  
 (~ 1200 astronomers)

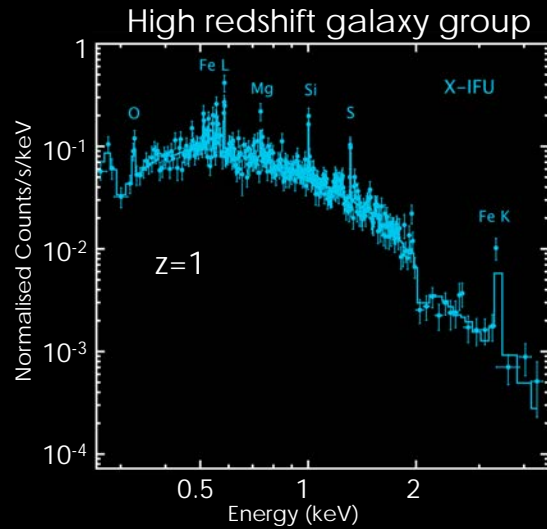
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**More information, white paper, 15 supporting papers at:**  
<http://the-athena-x-ray-observatory.eu/>

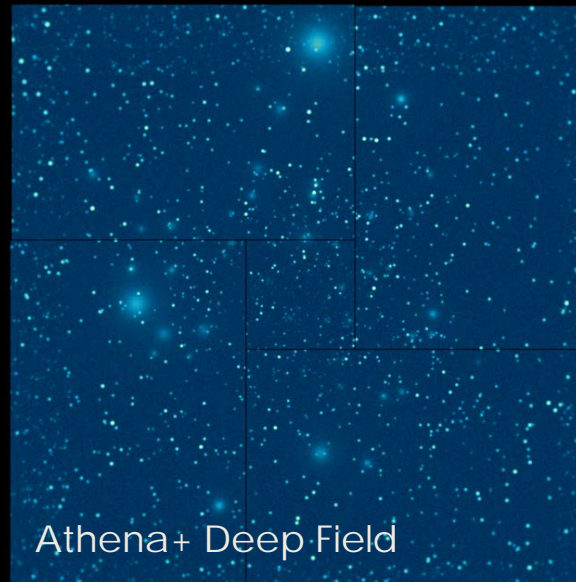
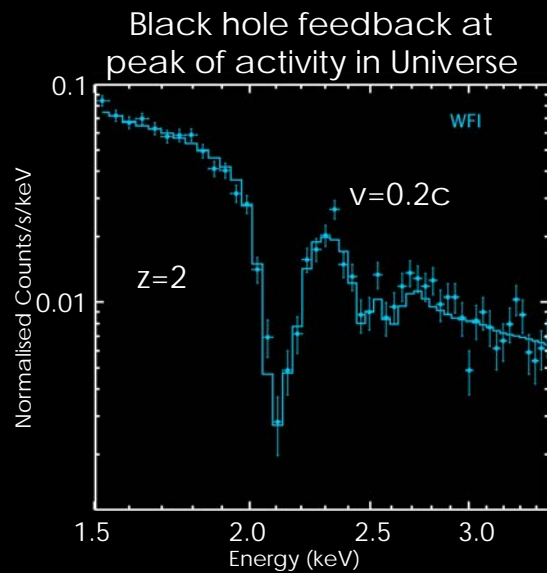
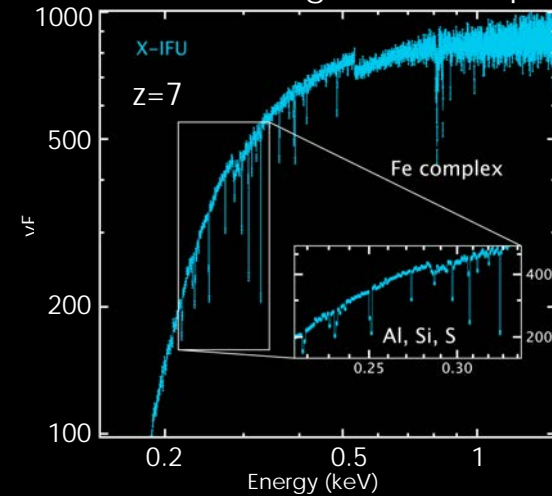
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# The first Deep Universe X-ray Observatory



Primordial stellar populations  
via GRB afterglow follow up



Nandra, Barret, Barcons, Fabian,  
den Herder, Piro, Watson et al.  
2013 arXiv 1306.2307

Obscured black hole in  
the early Universe

